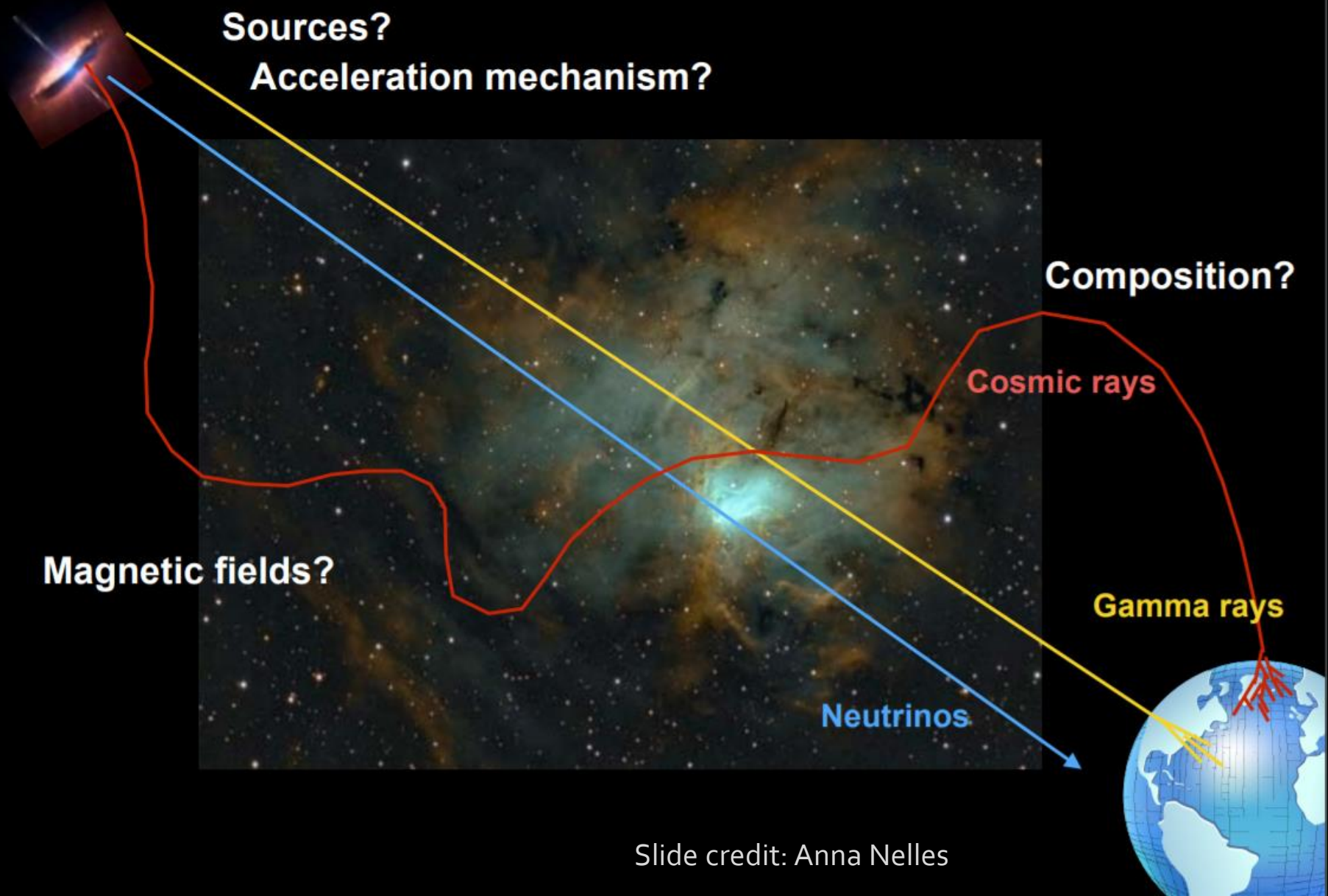


# Direct detection of Cosmic Rays: Techniques, challenges and early results

Ryan Monroe, Anna Nelles, Gregg Hallinan,  
Michael Eastwood, Marin Anderson,  
and many others!

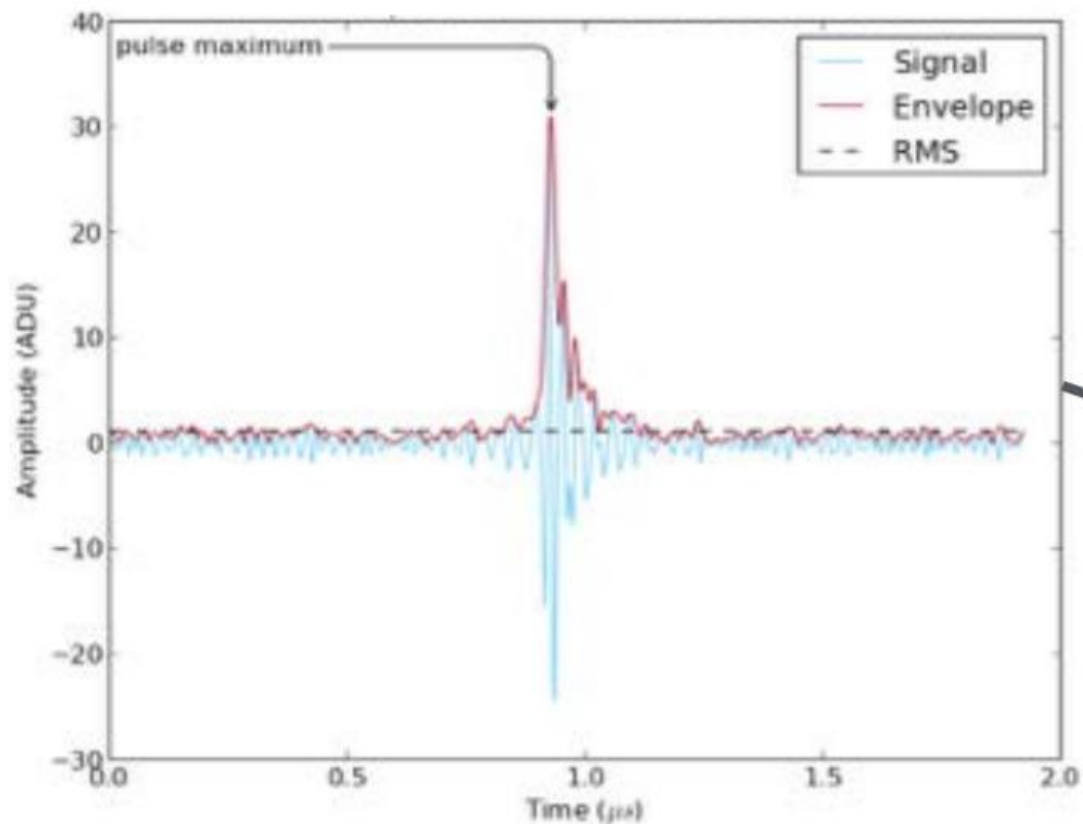
# Astroparticle Physics

At the highest energies

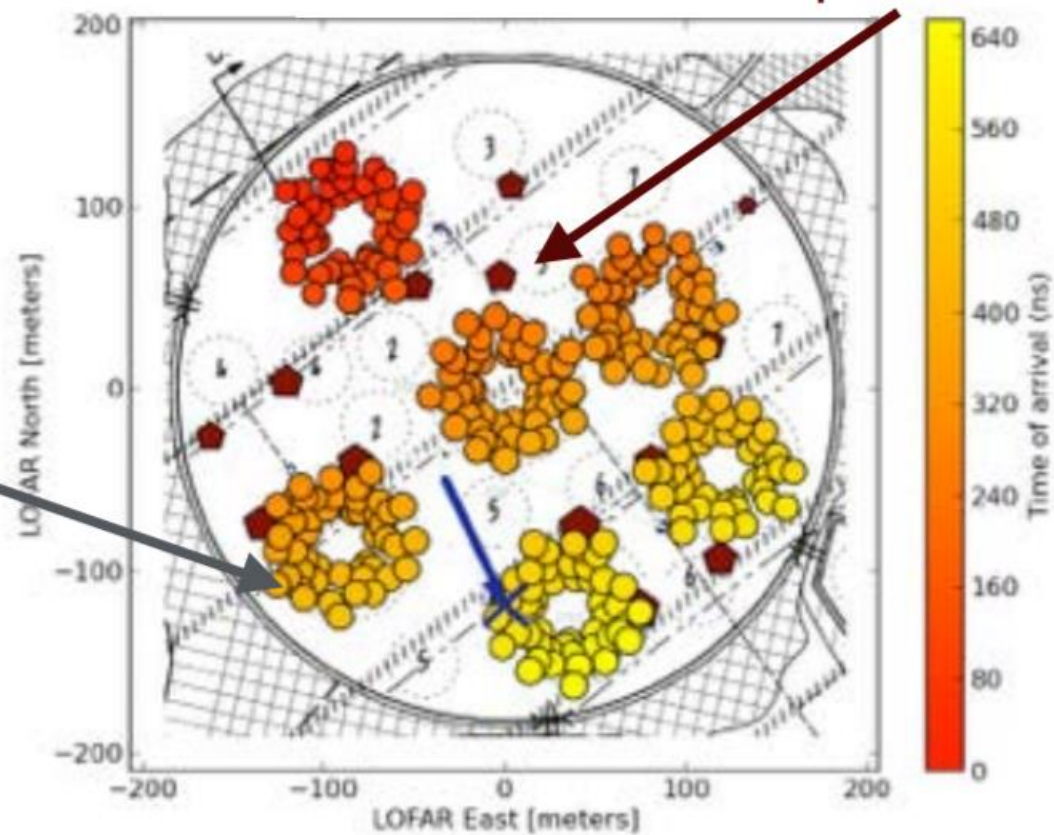


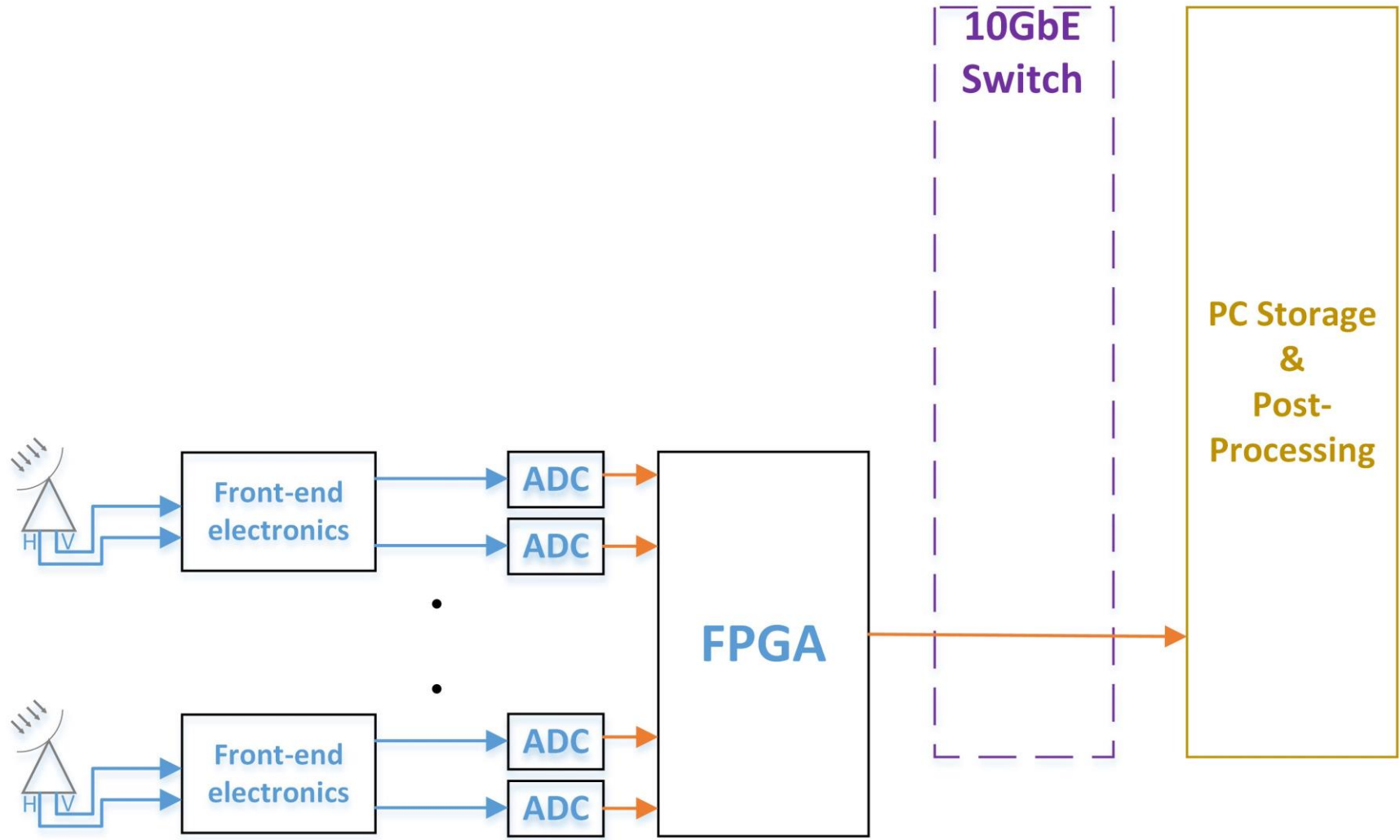
Slide credit: Anna Nelles

## Single antenna data LOFAR 30-80 MHz



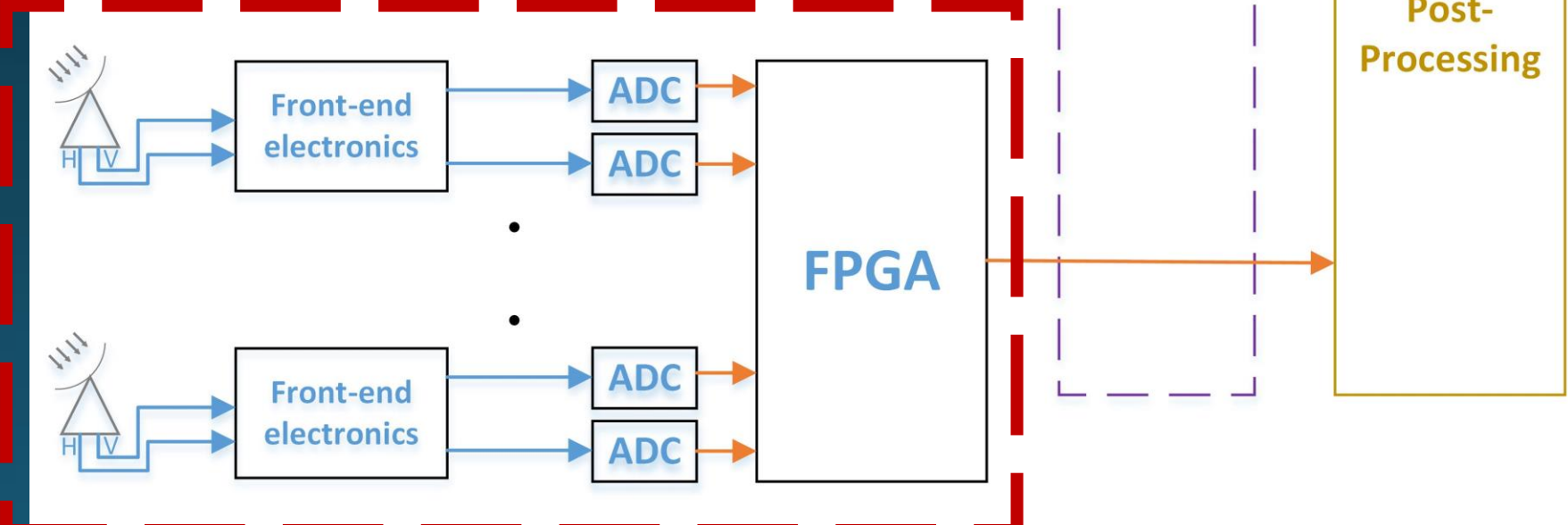
Particle detectors provide trigger



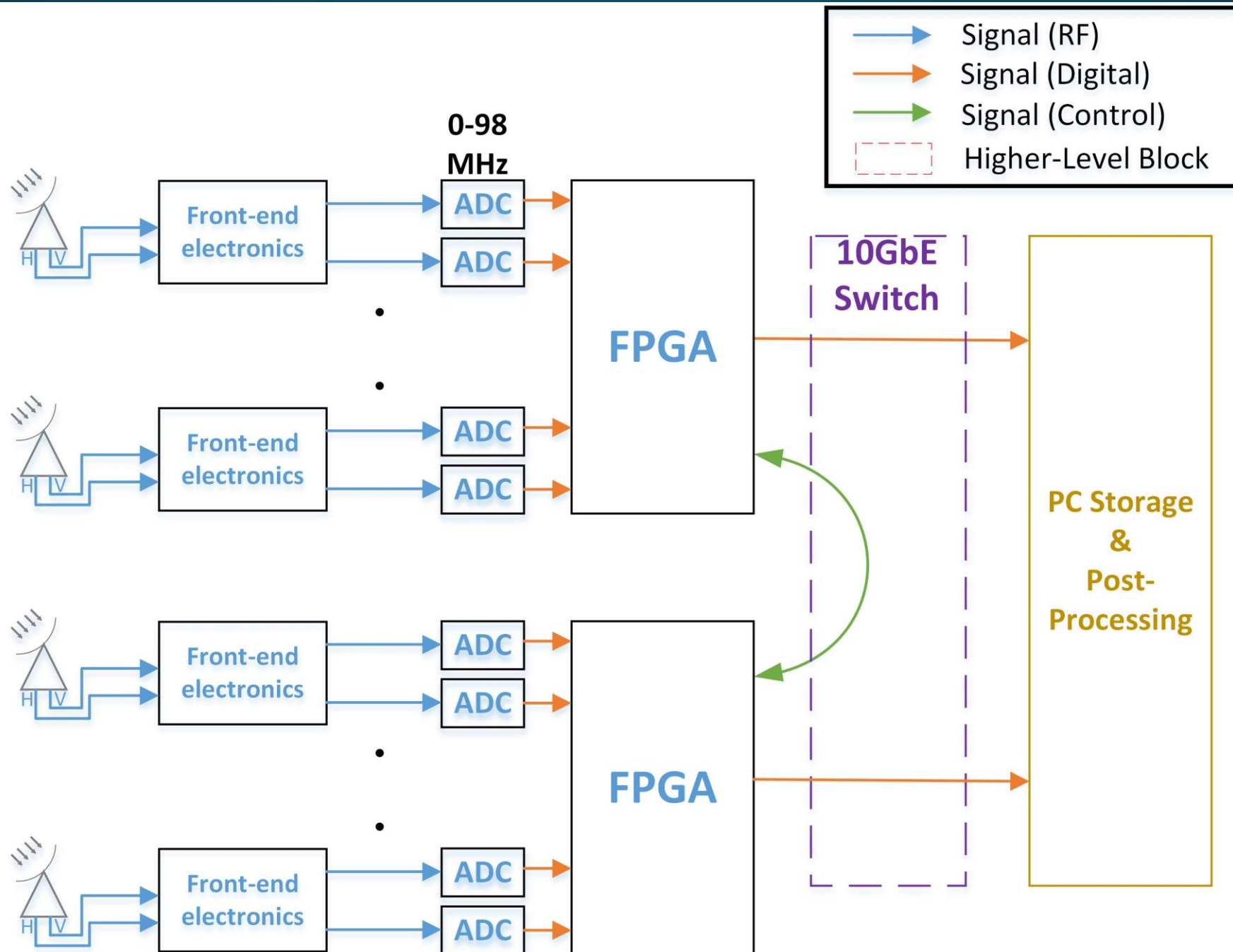




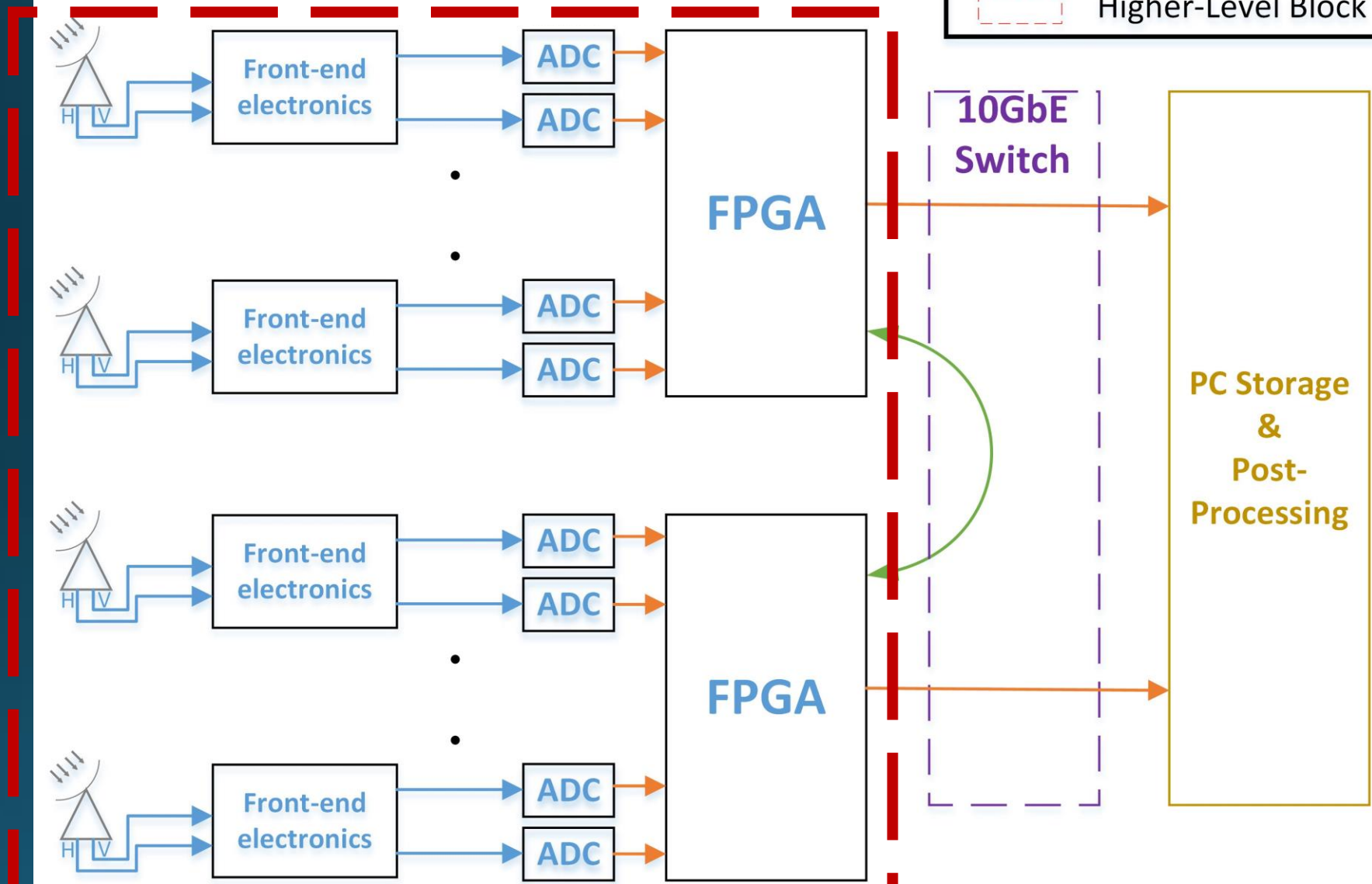
**16 Antennas  
(32 inputs) per FPGA**



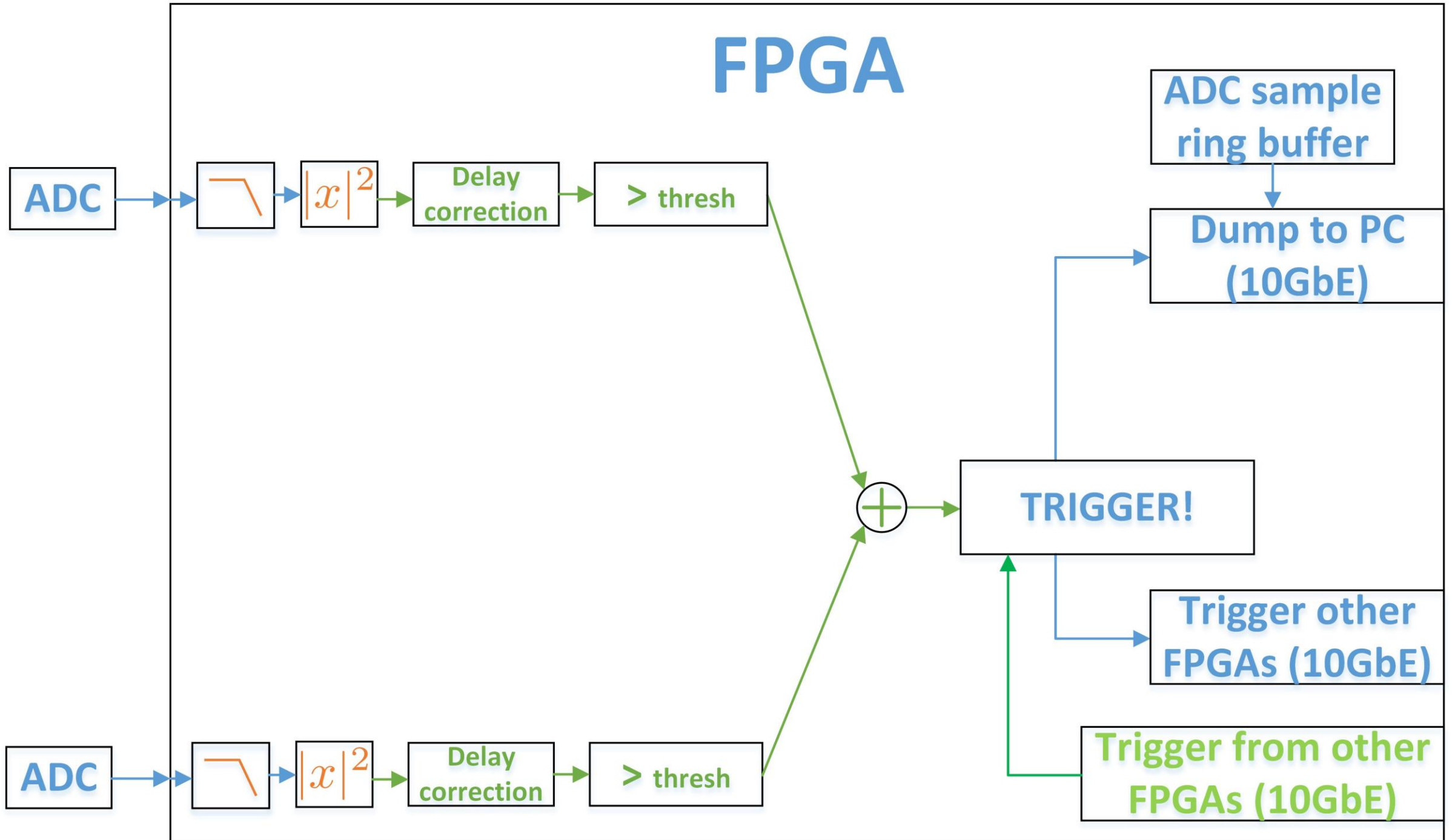




# 16 FPGAs total (512 inputs total)



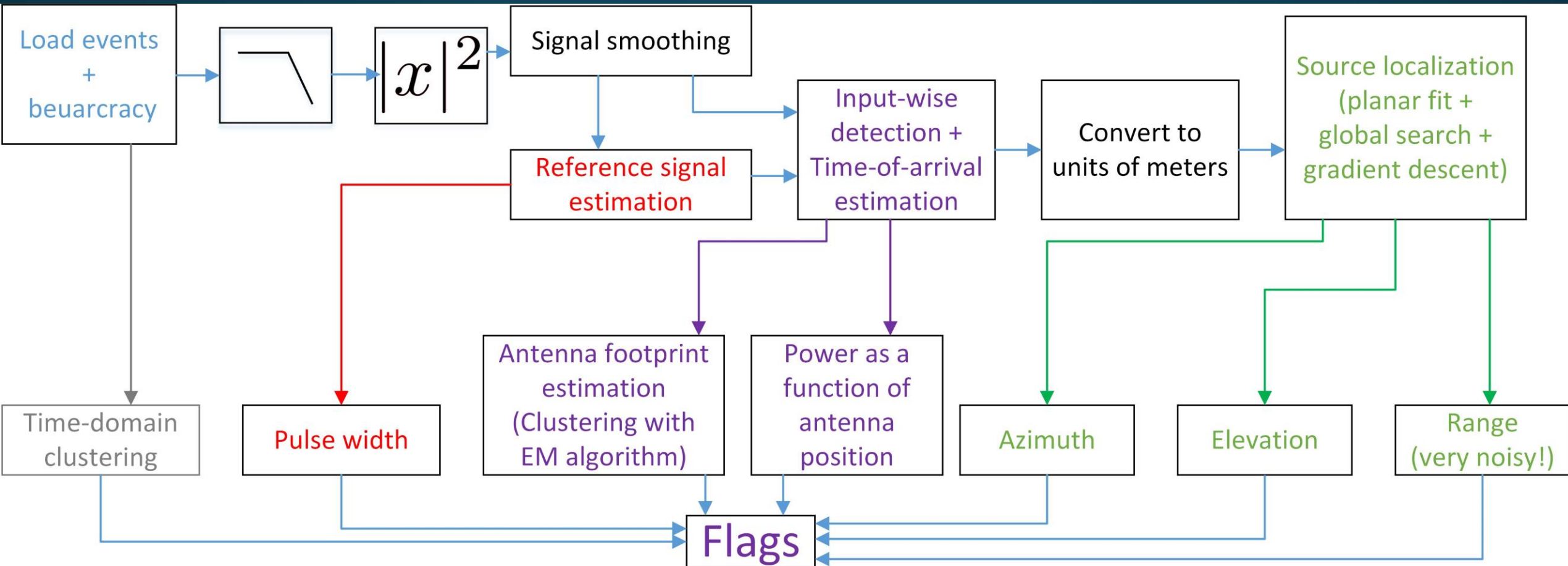
# FPGA

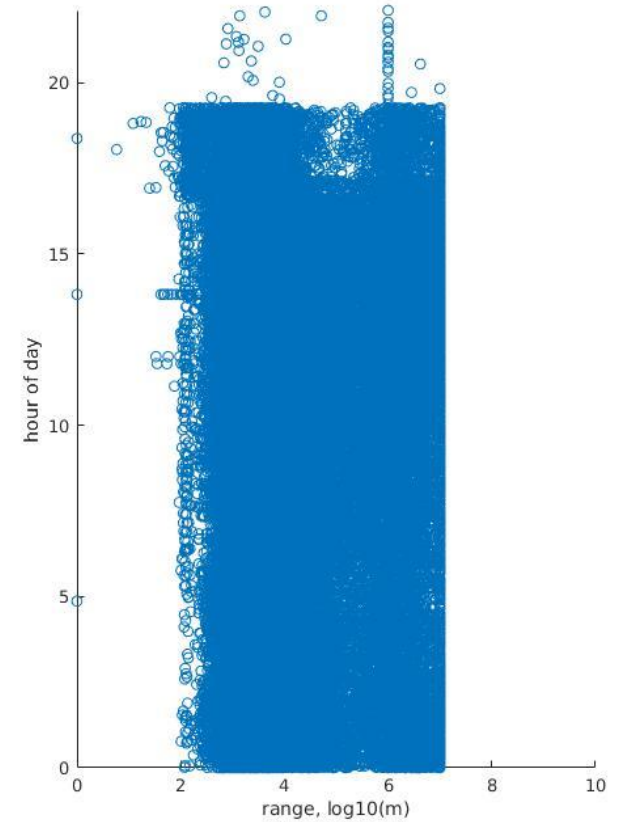
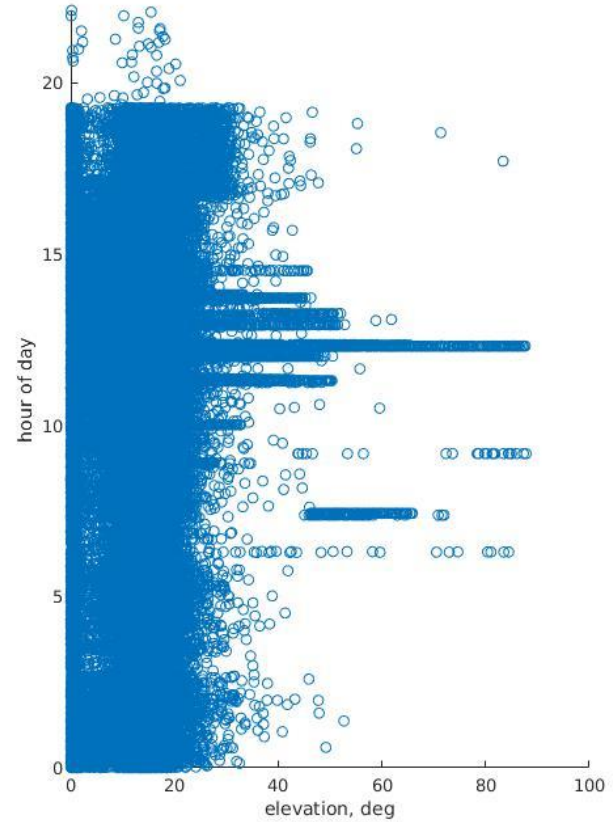
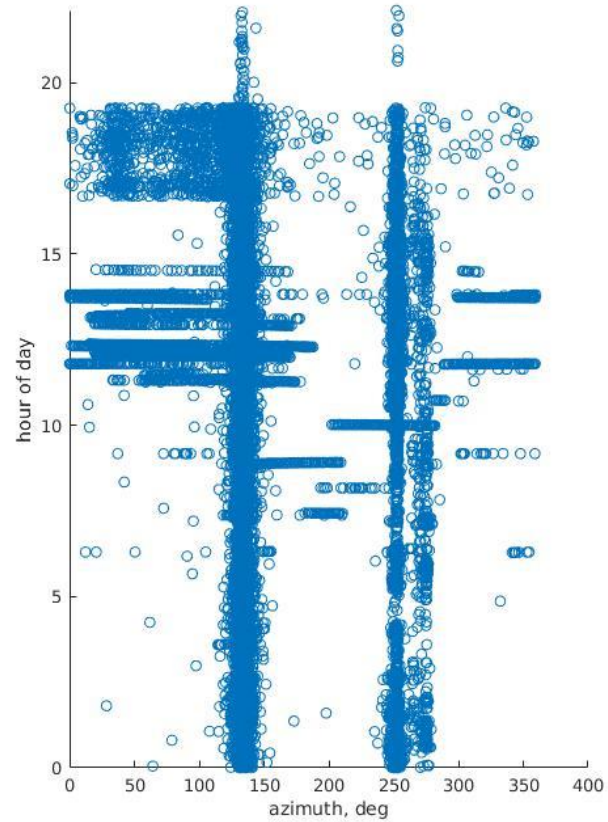
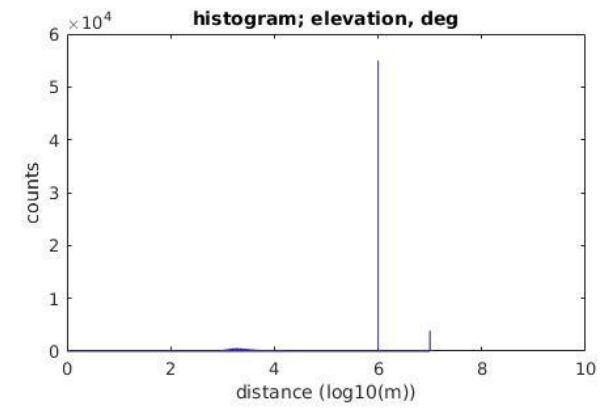
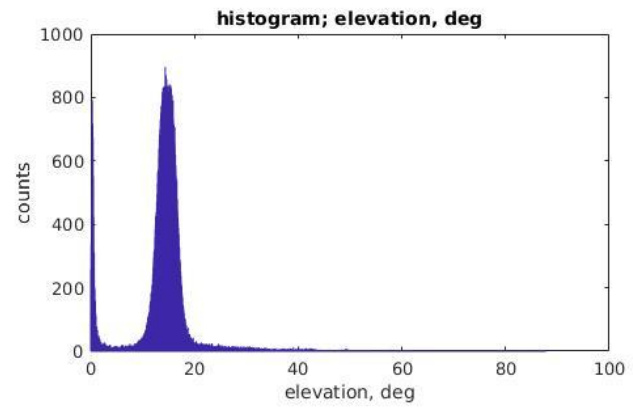
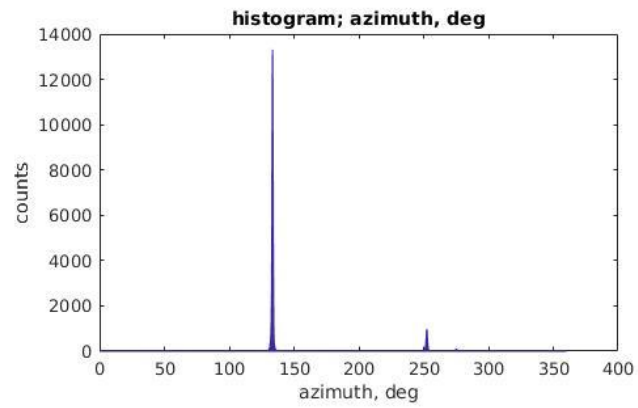


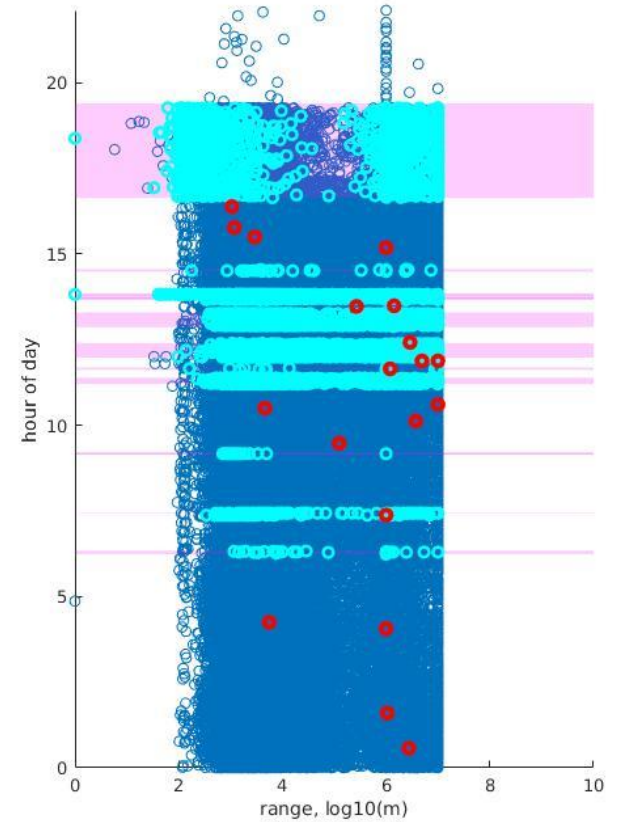
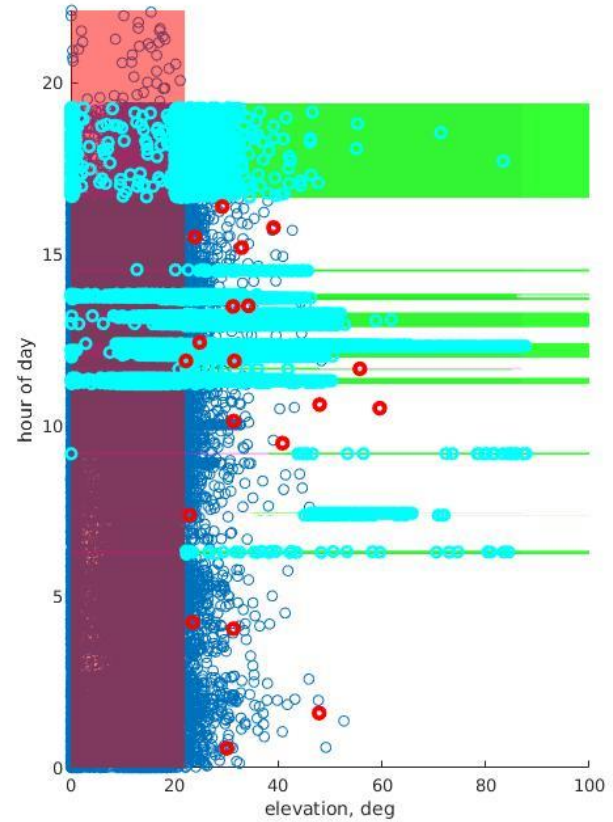
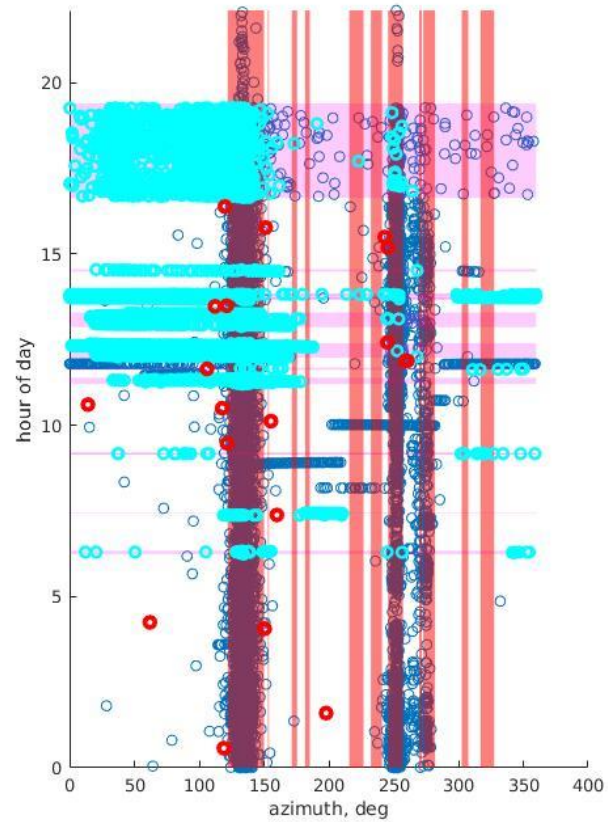
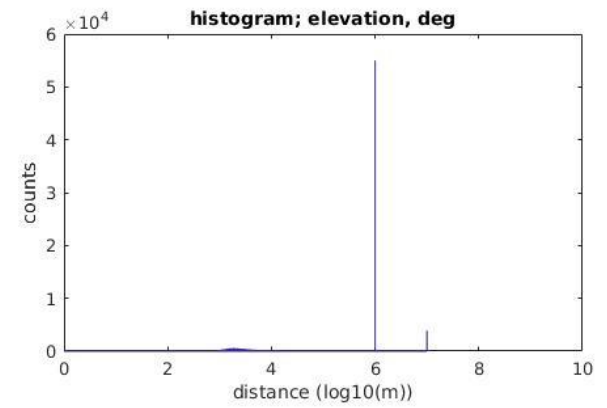
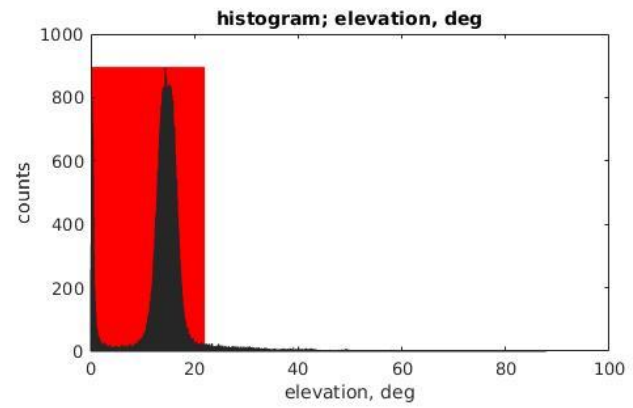
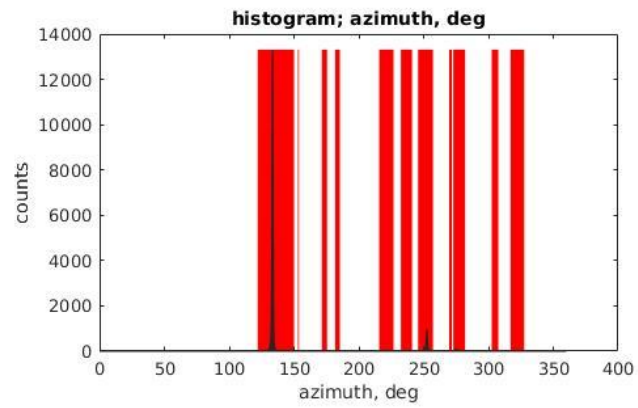


# PC Post-Processing

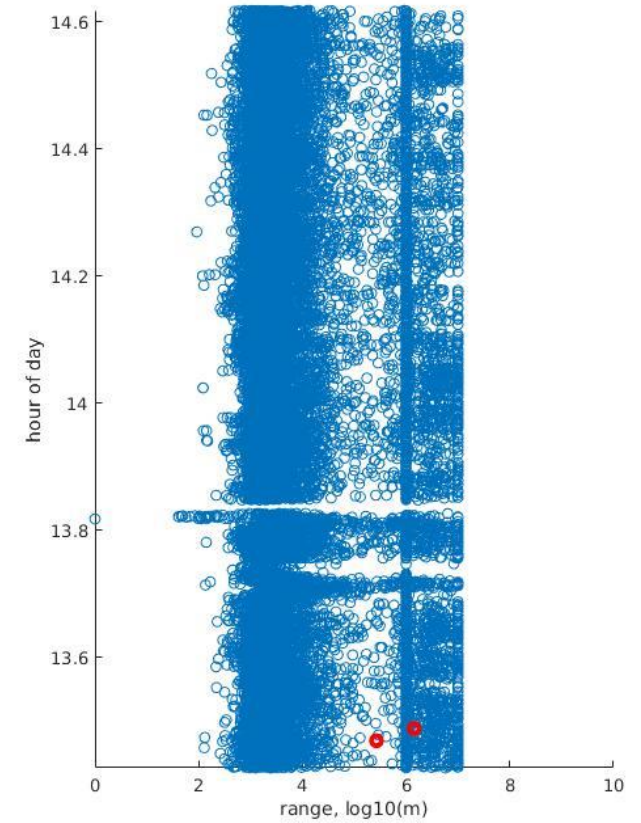
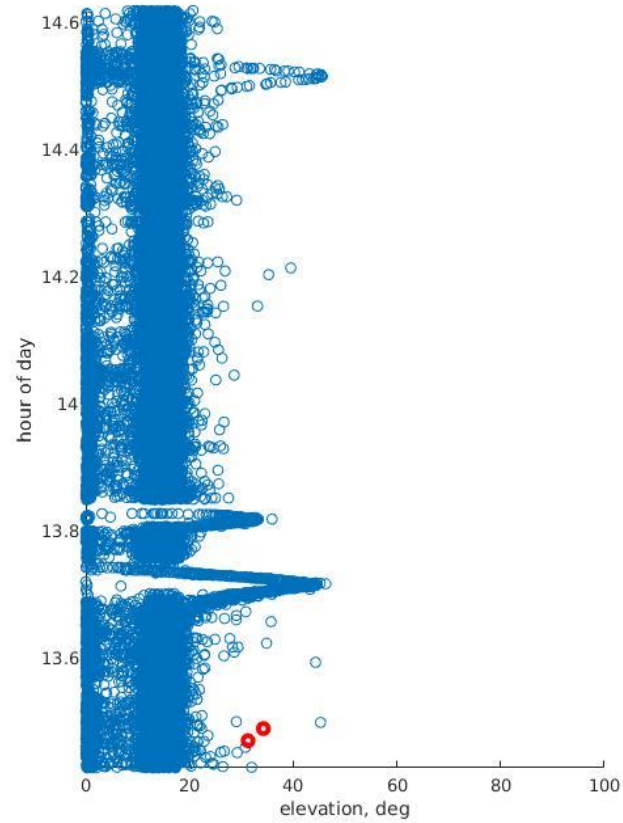
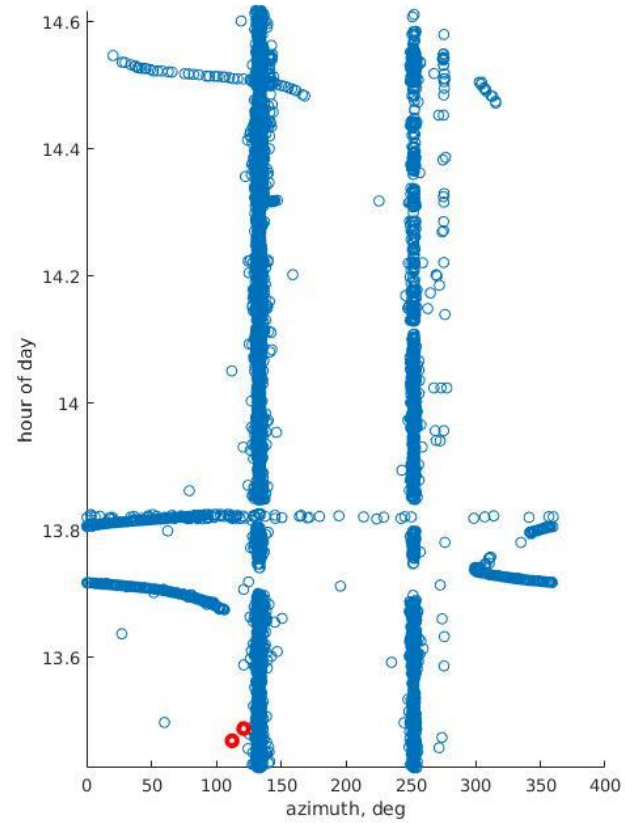
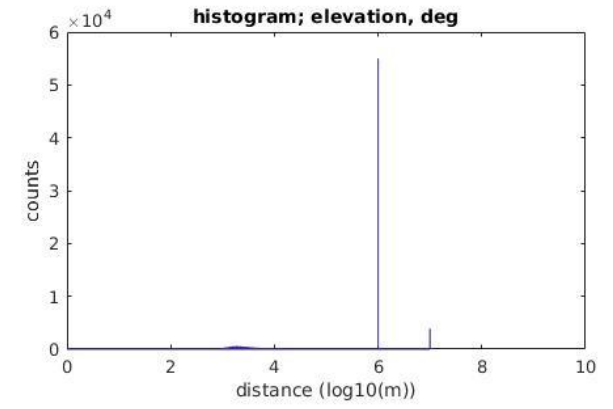
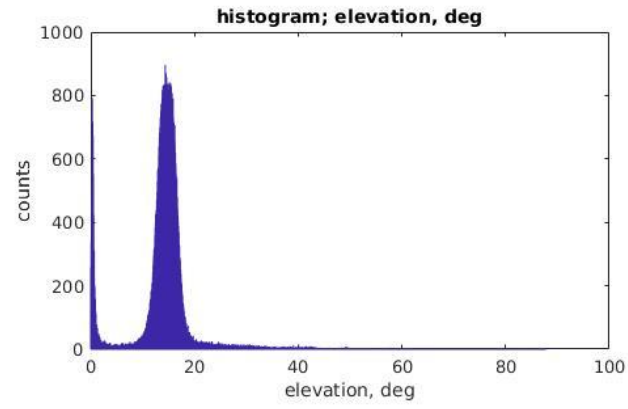
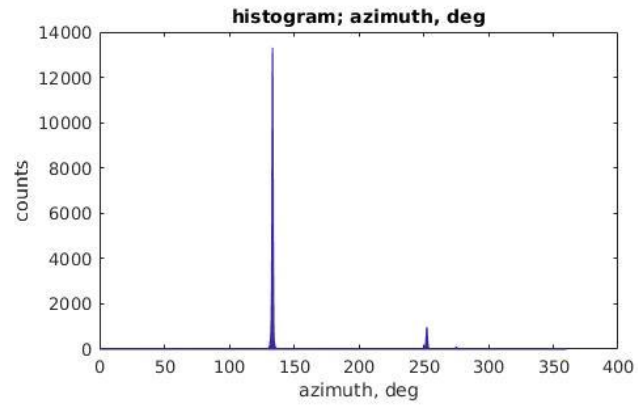
Challenge: No strong priors on CRs until we have successful detections



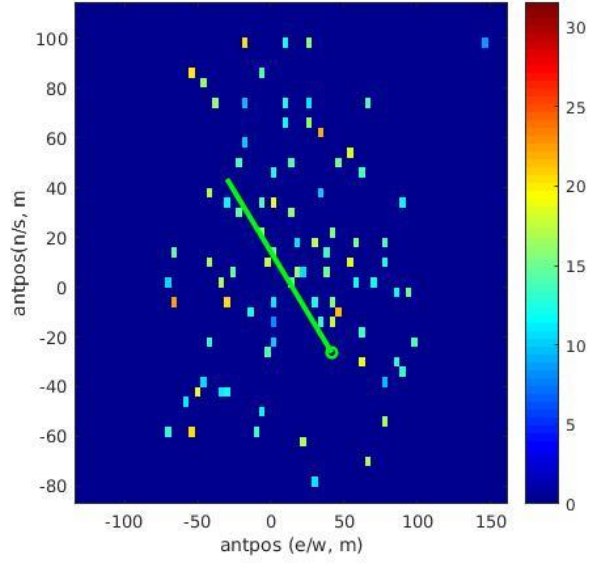




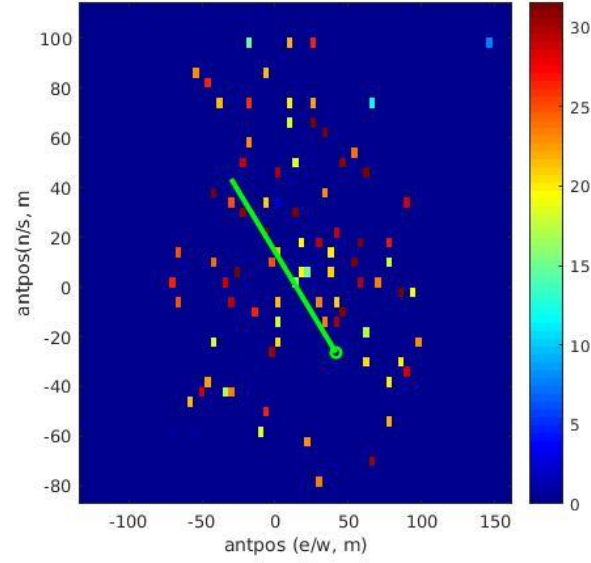




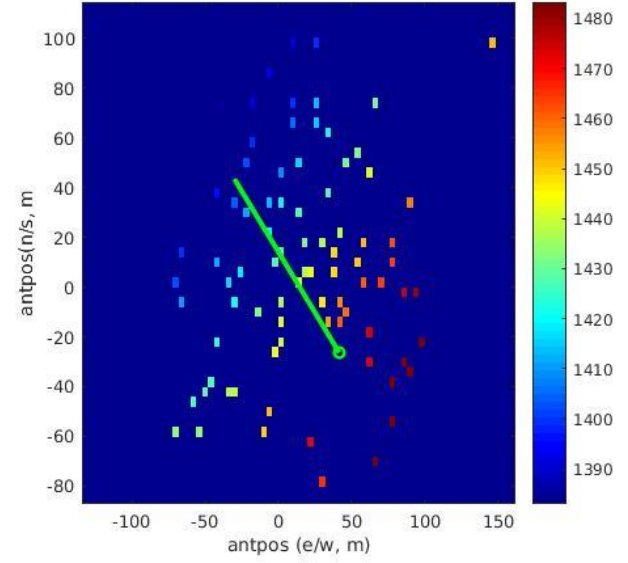
Candidate CR event: relative arrival pows (V pol)



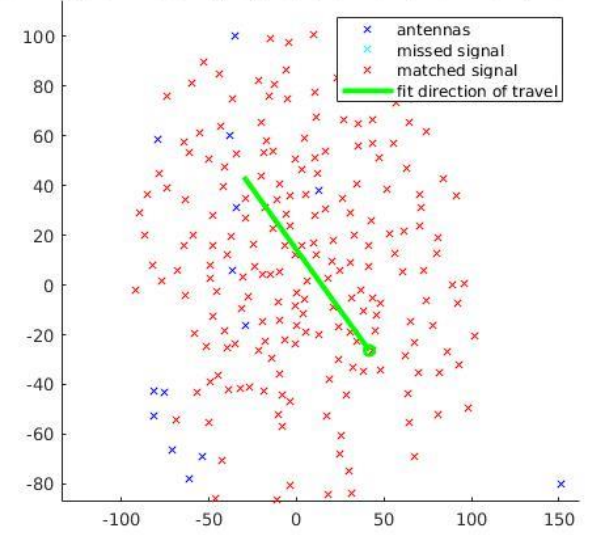
Candidate CR event: relative arrival pows (H pol)



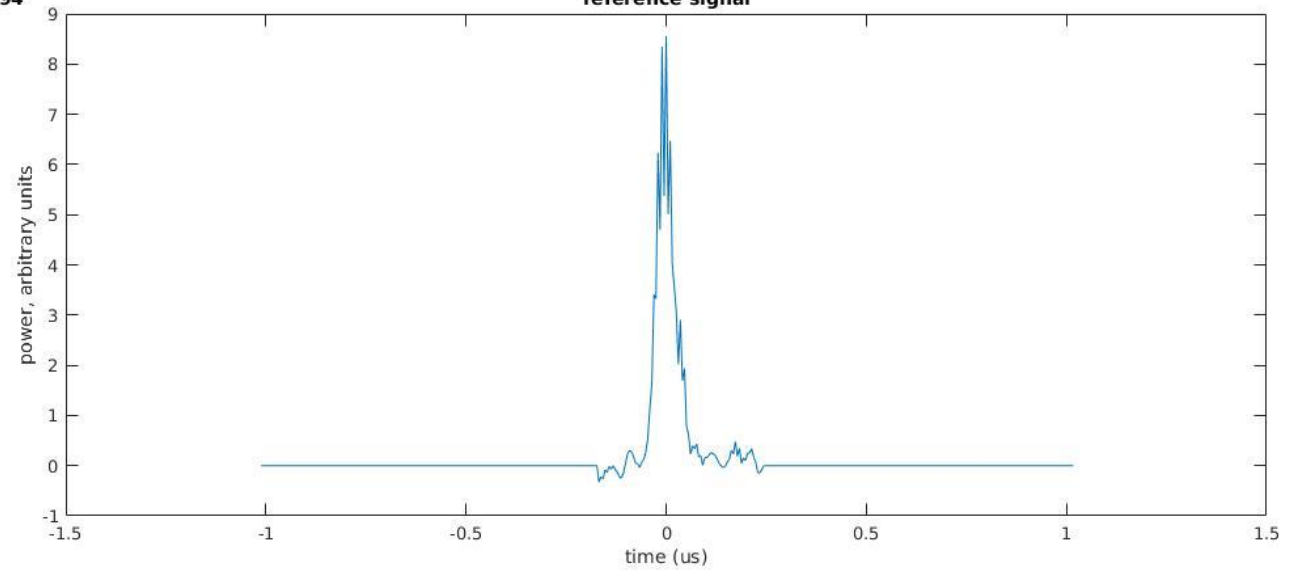
Candidate CR event: arrival times



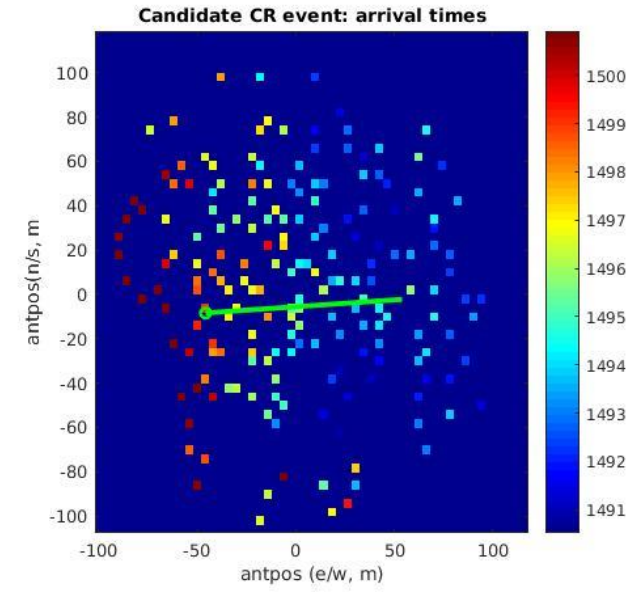
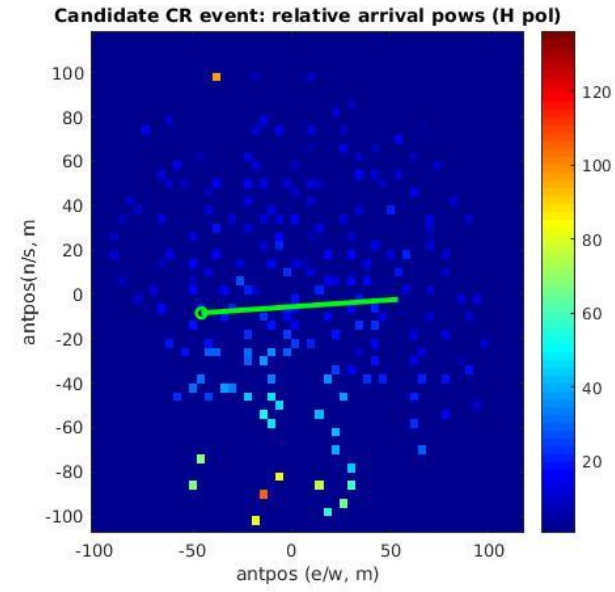
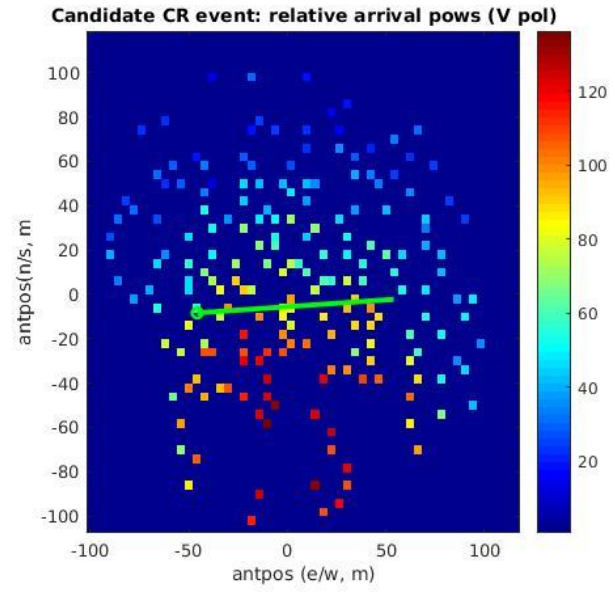
az: 135.7446, | el: -18.521, | log10(r): 3.7491 | nOK: 292 | indThis: 133294



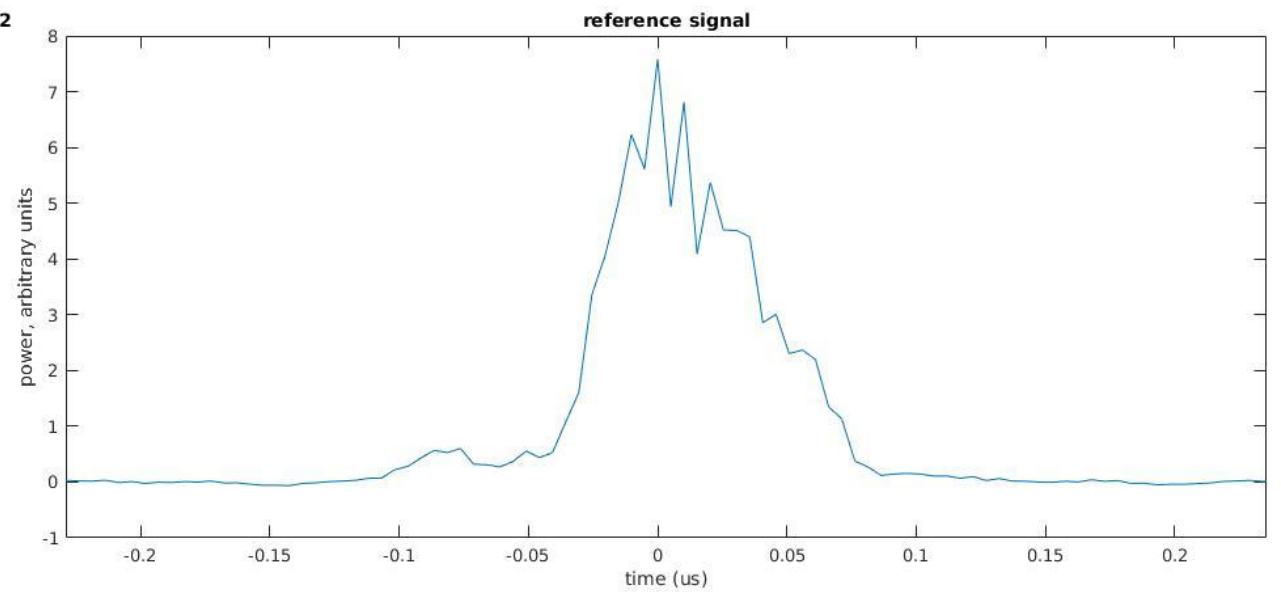
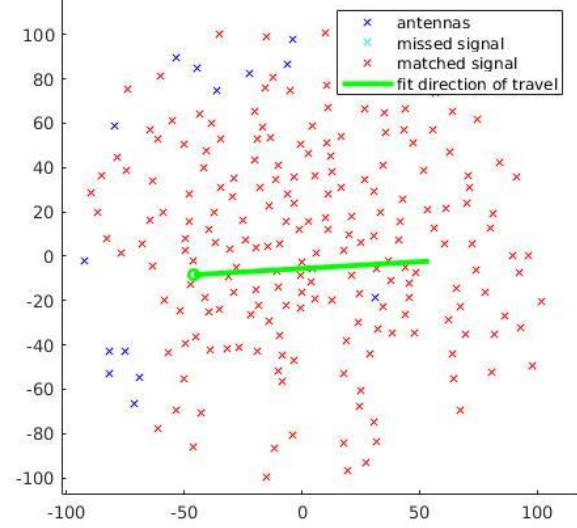
reference signal

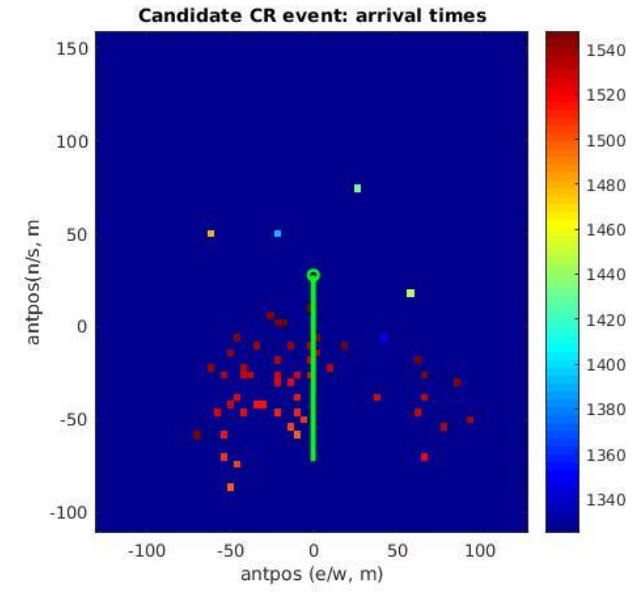
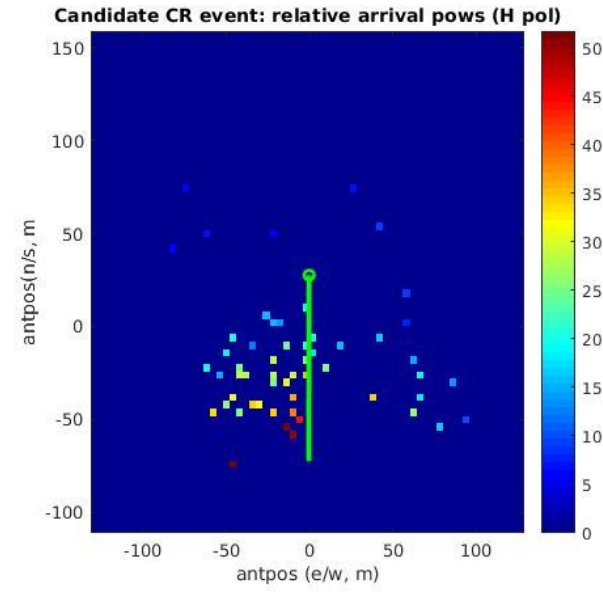
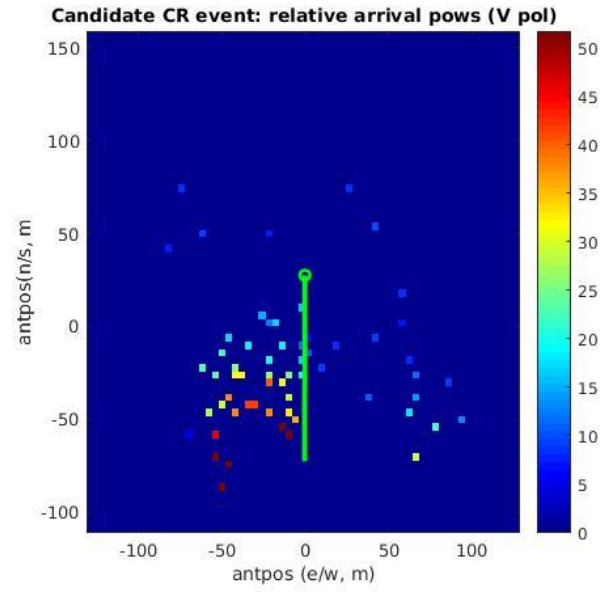




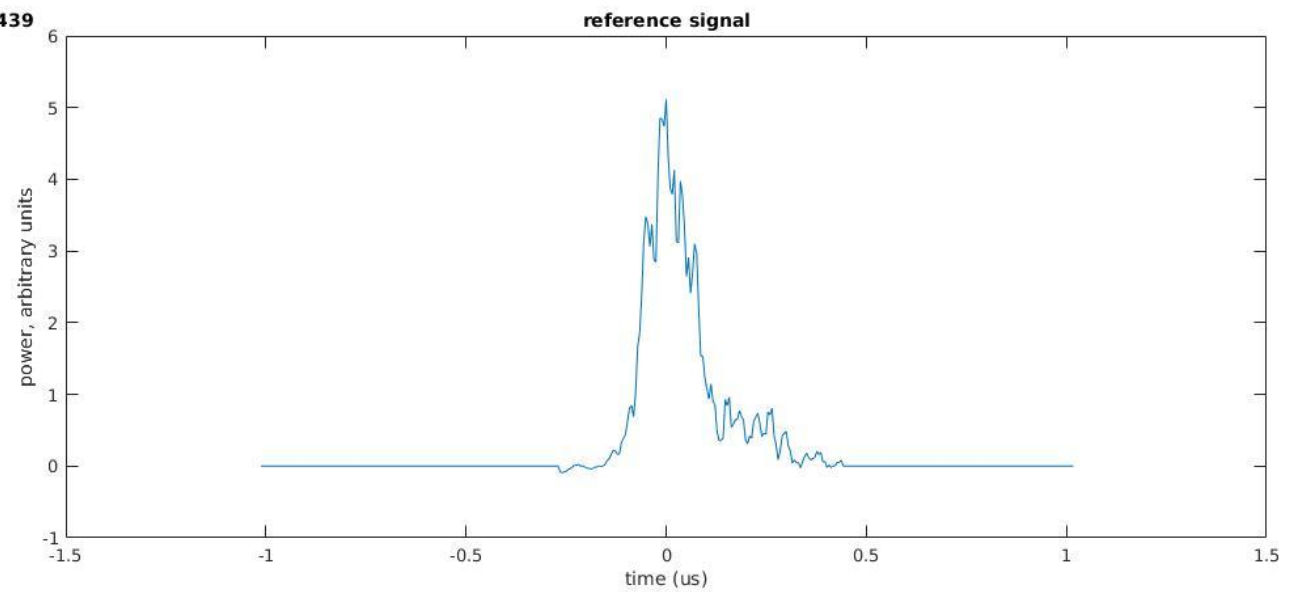
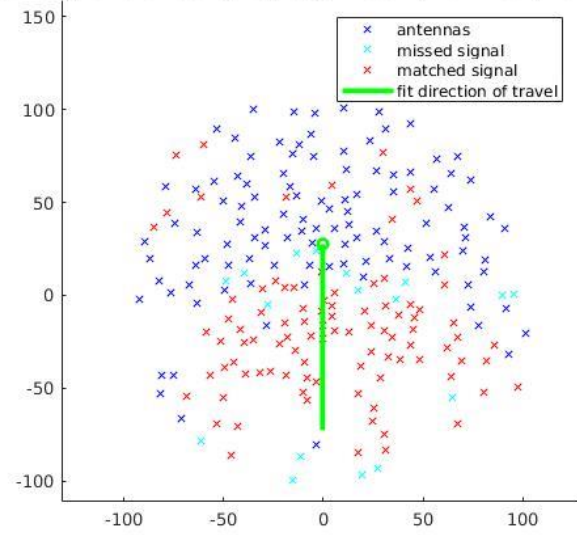


az: 3.524, | el: -85.2072, | log10(r): 3.4381 | nOK: 301 | indThis: 238482





az: 269.9183, | el: 0.024854, | log10(r): 2.0991 | nOK: 130 | indThis: 106439







Wilkerson

Peterson Mill

Keough  
Hot Springs

Round  
Mountain

Sugarloaf

OVRO-LWA

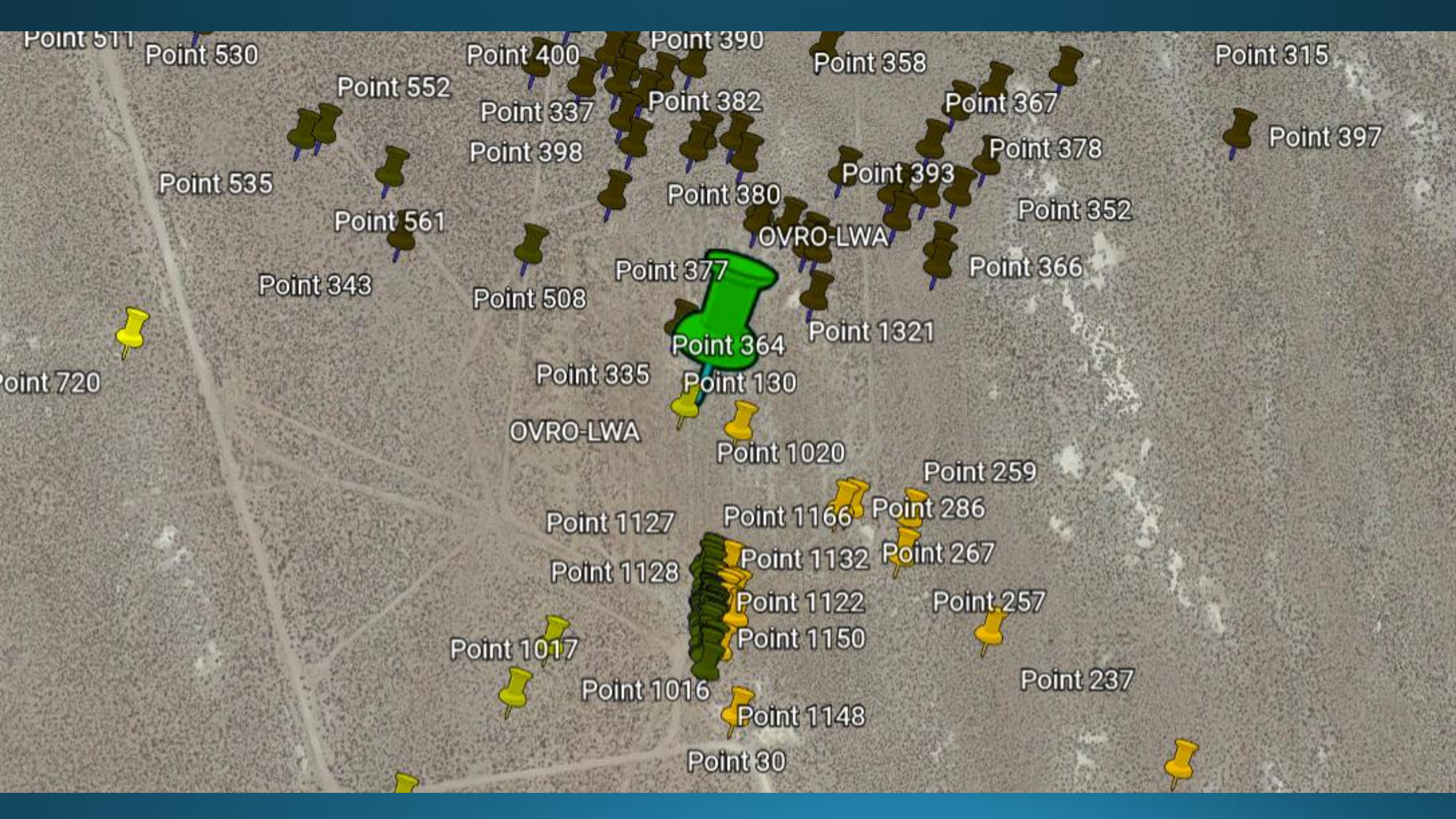
Zurich

Black  
Mountain

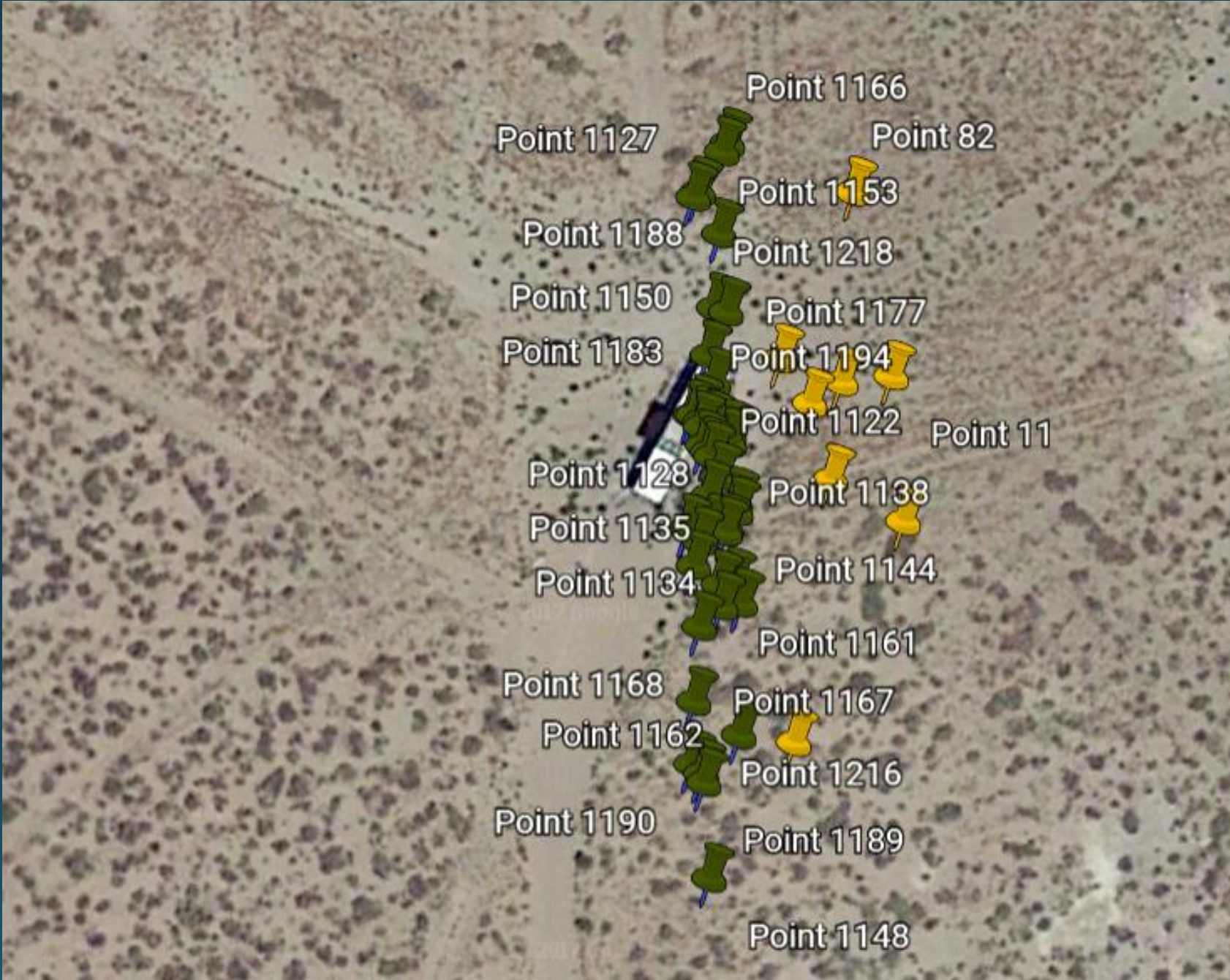


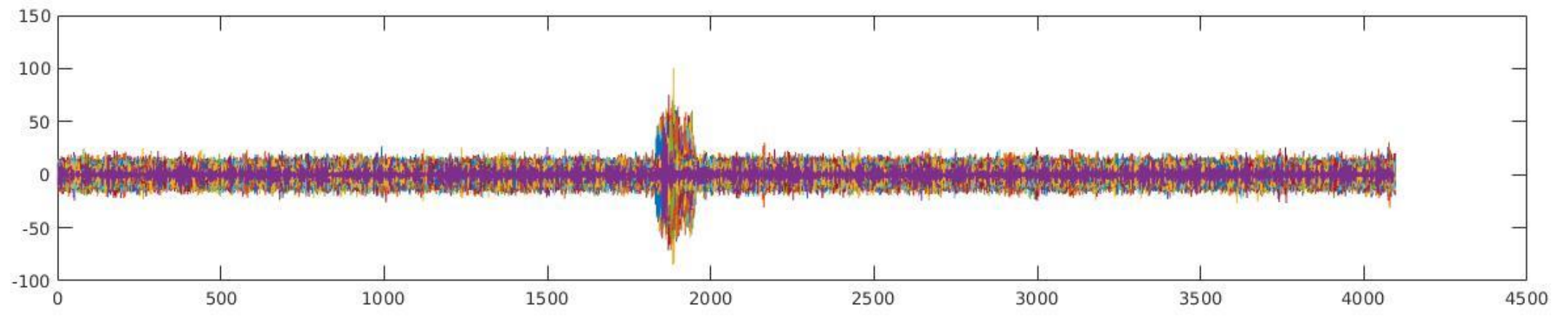


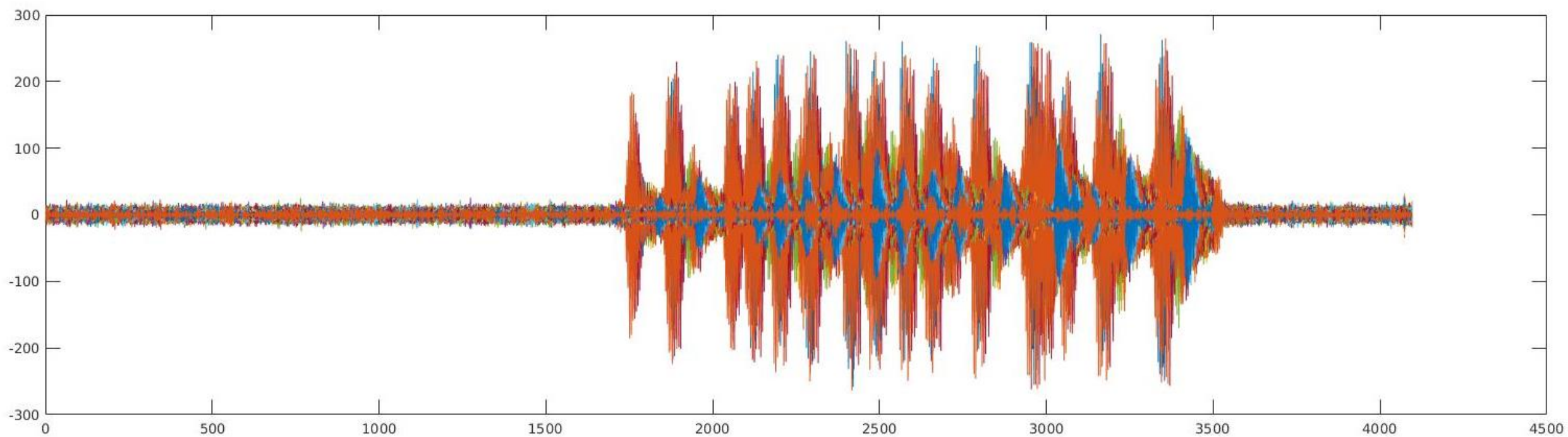




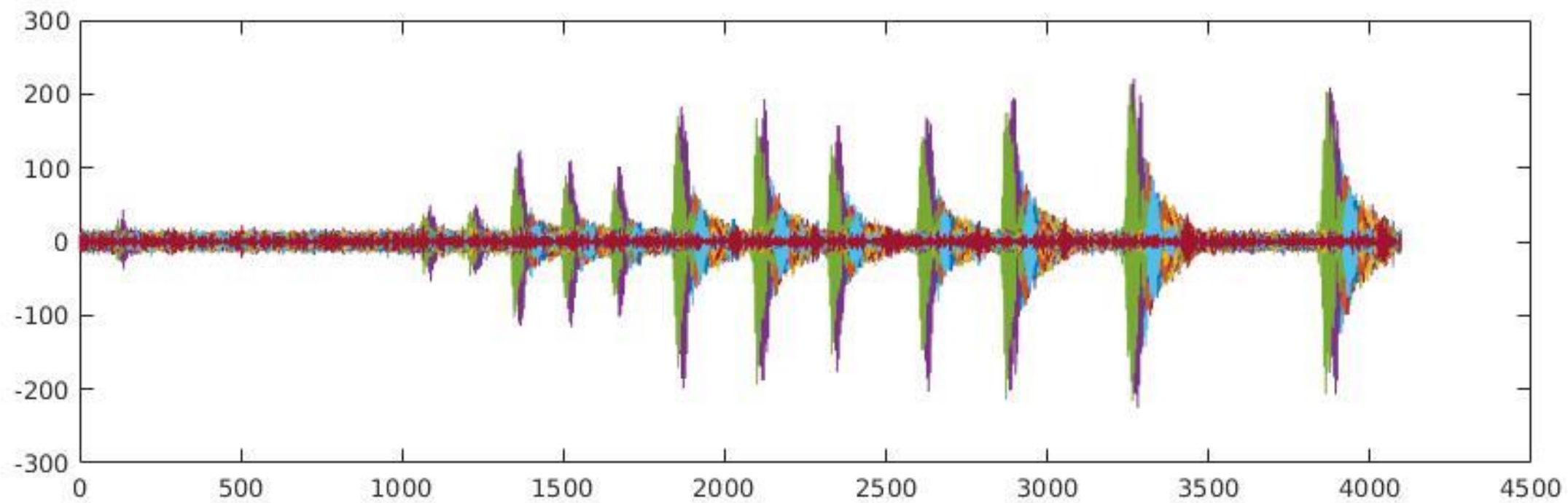


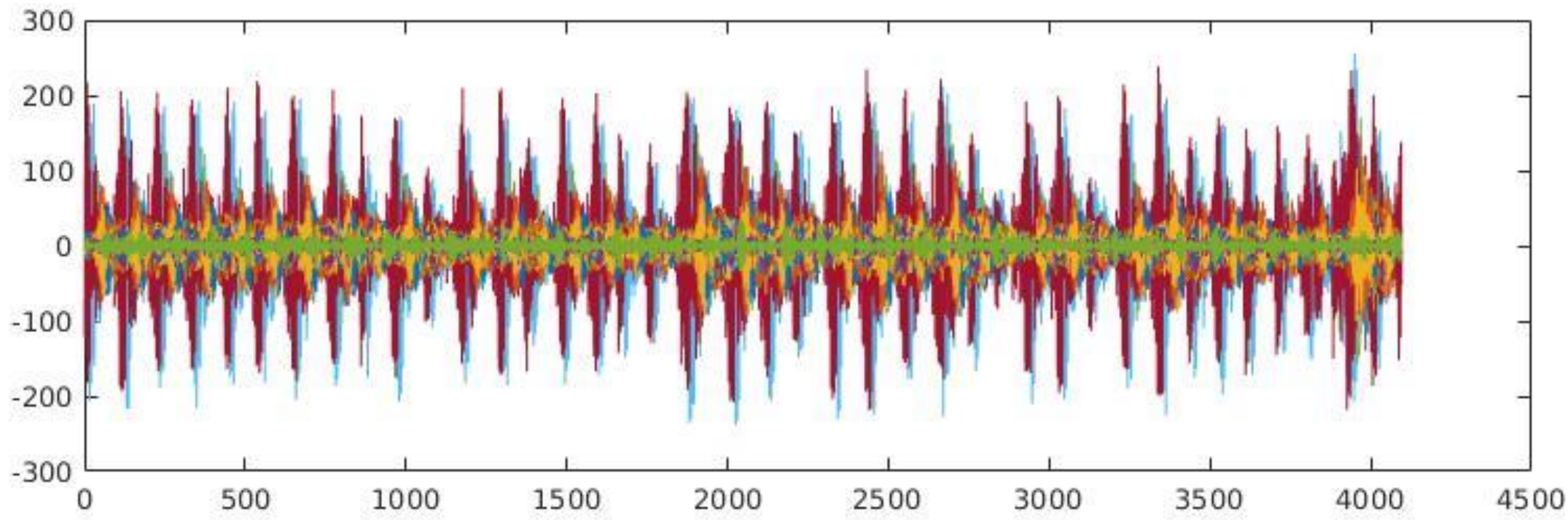




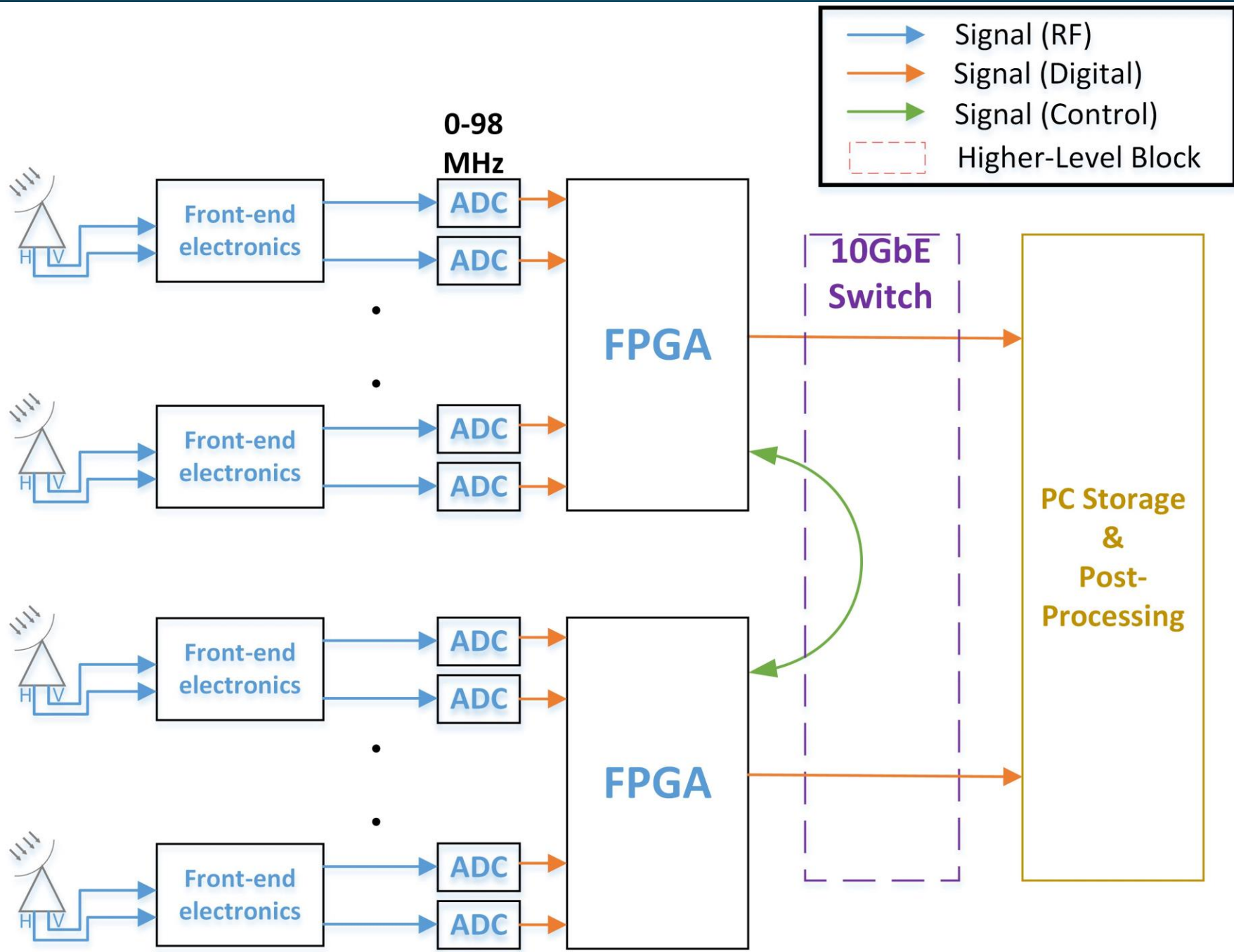


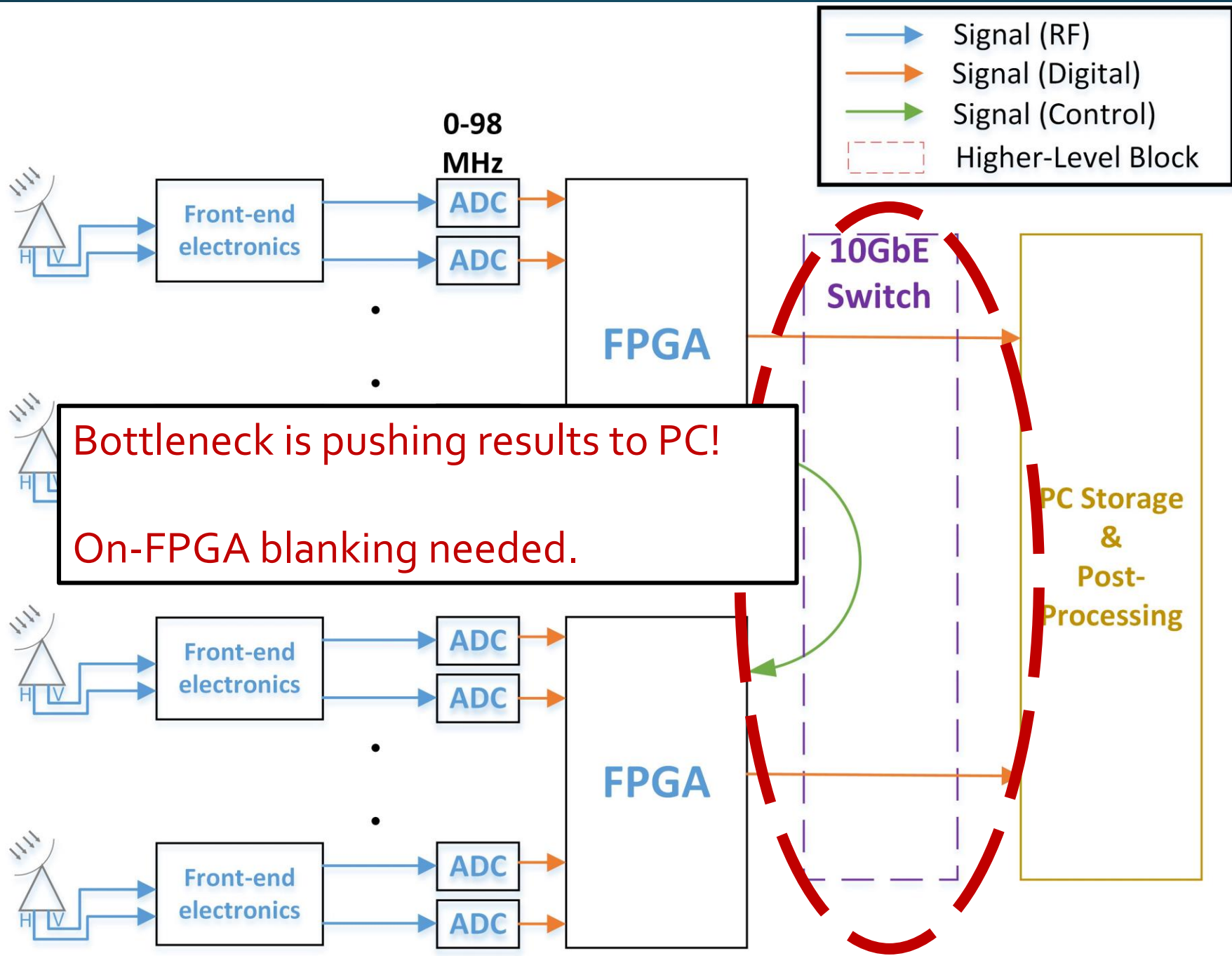




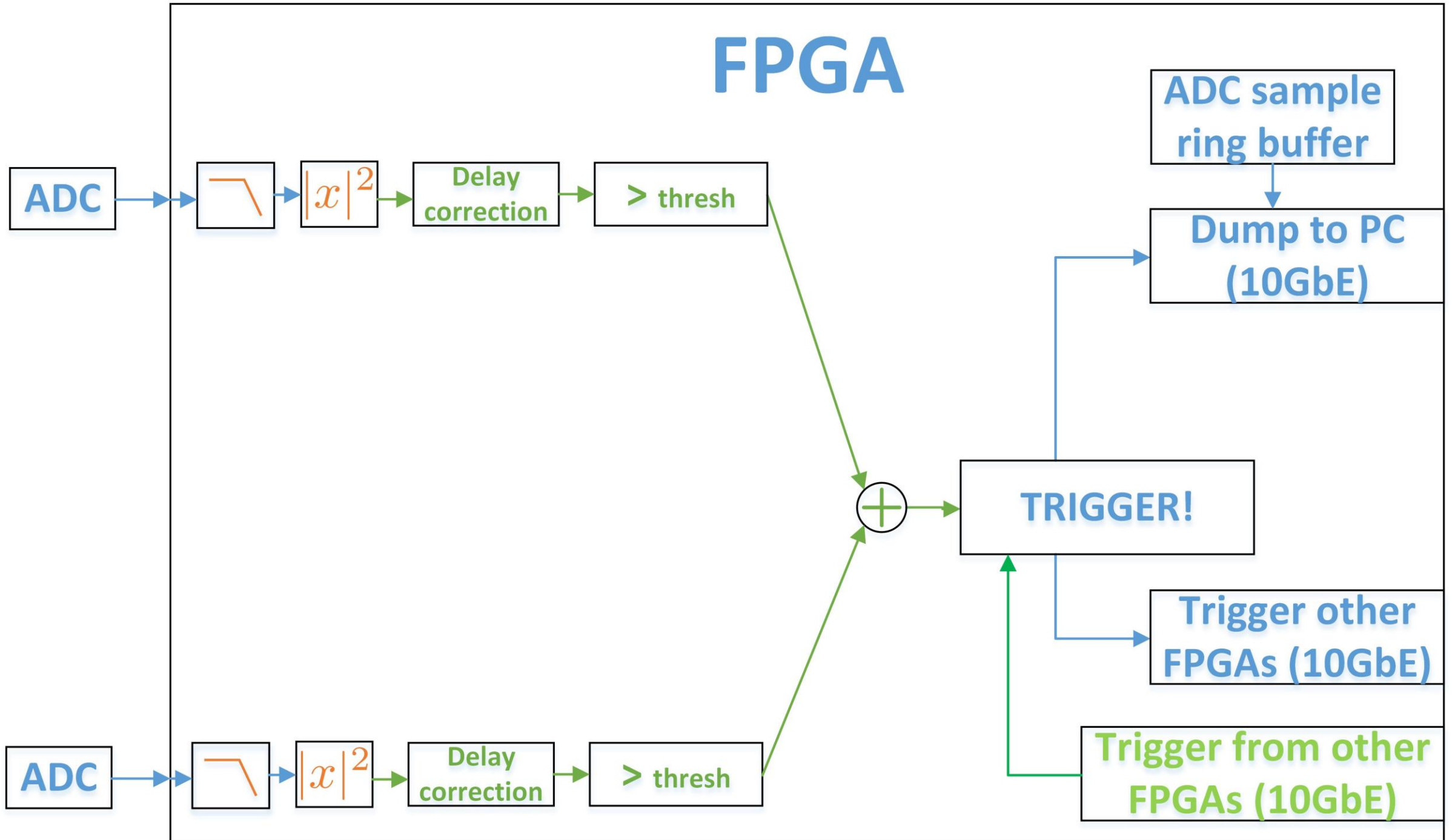




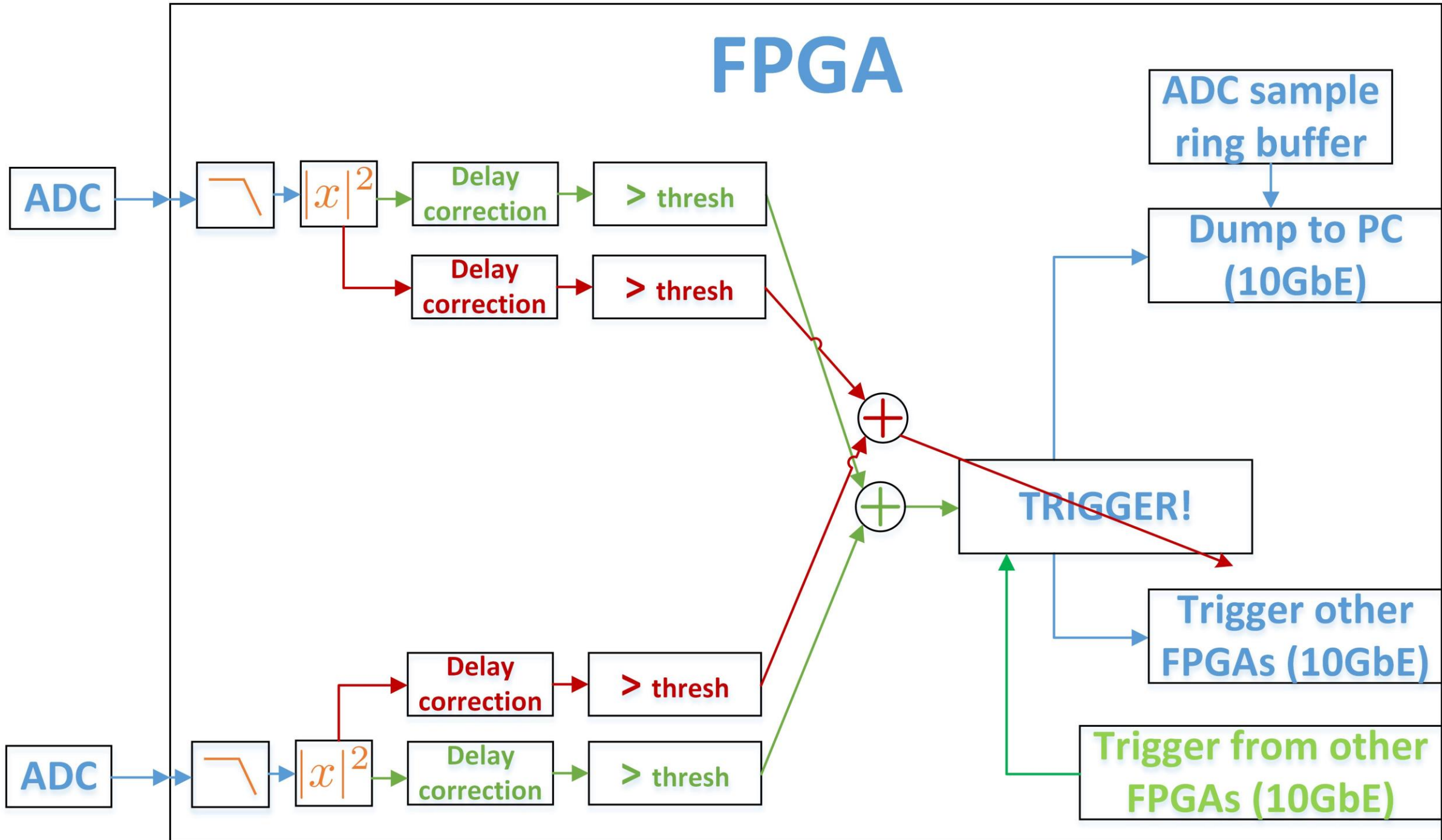




# FPGA

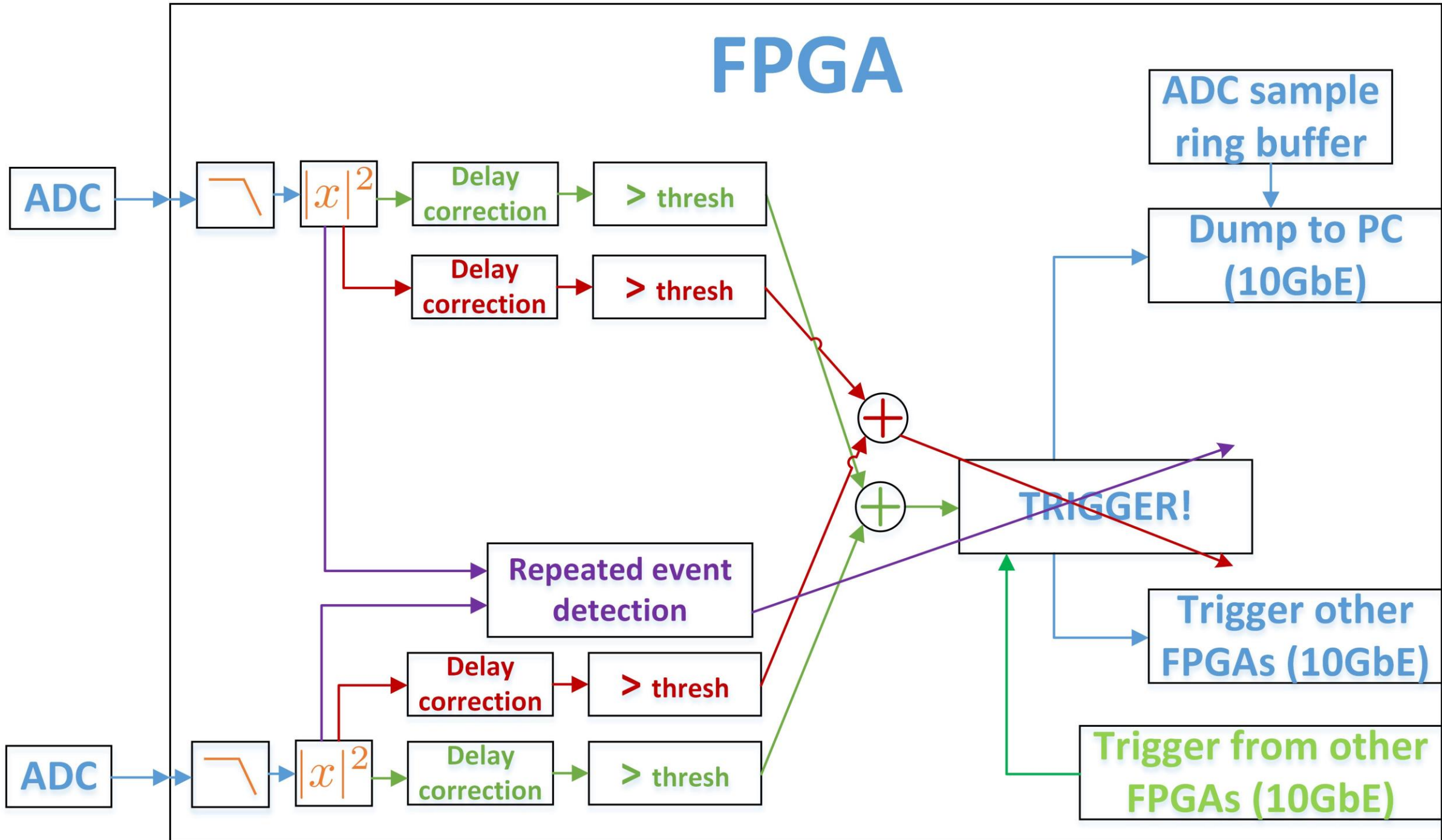


# FPGA

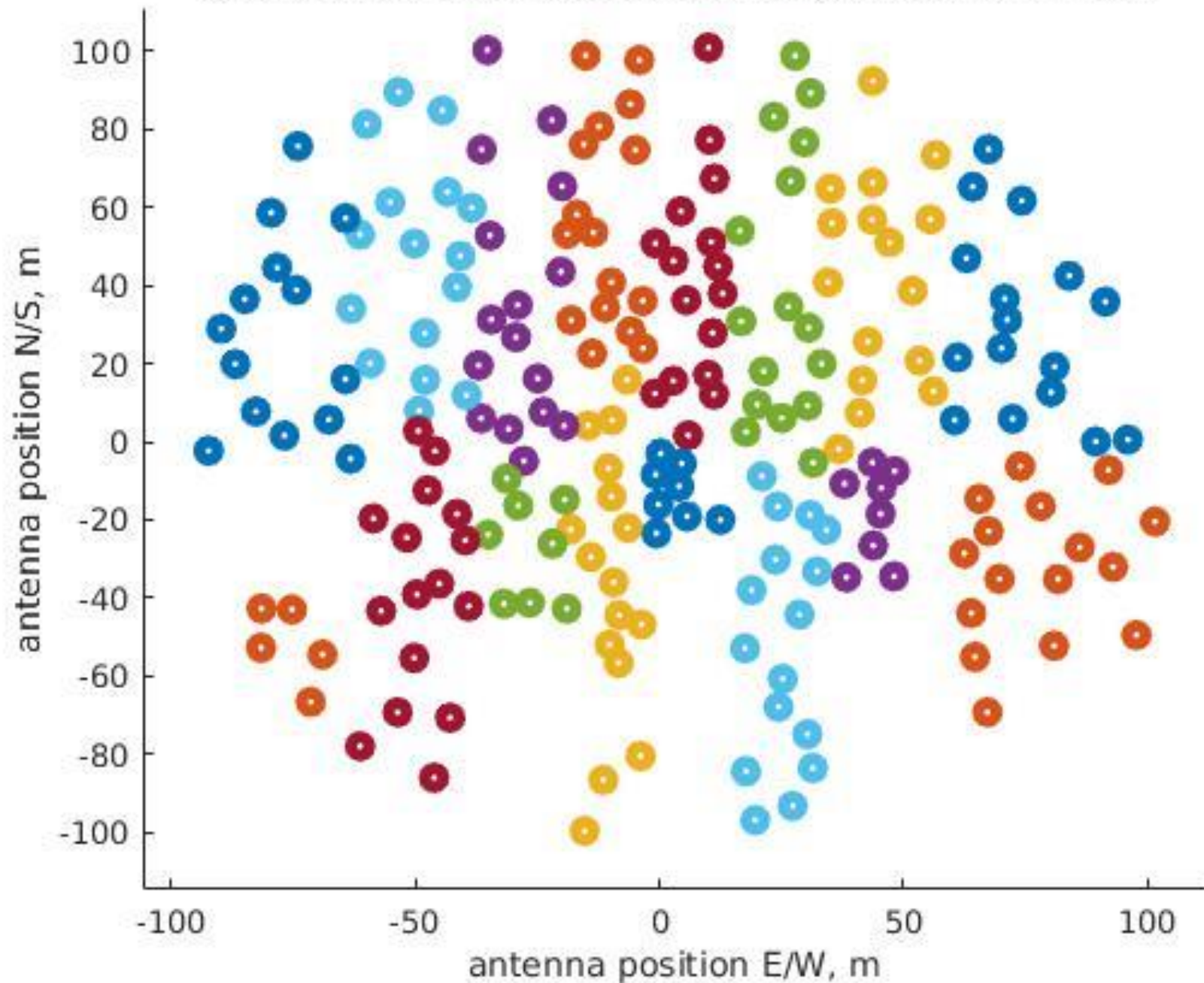


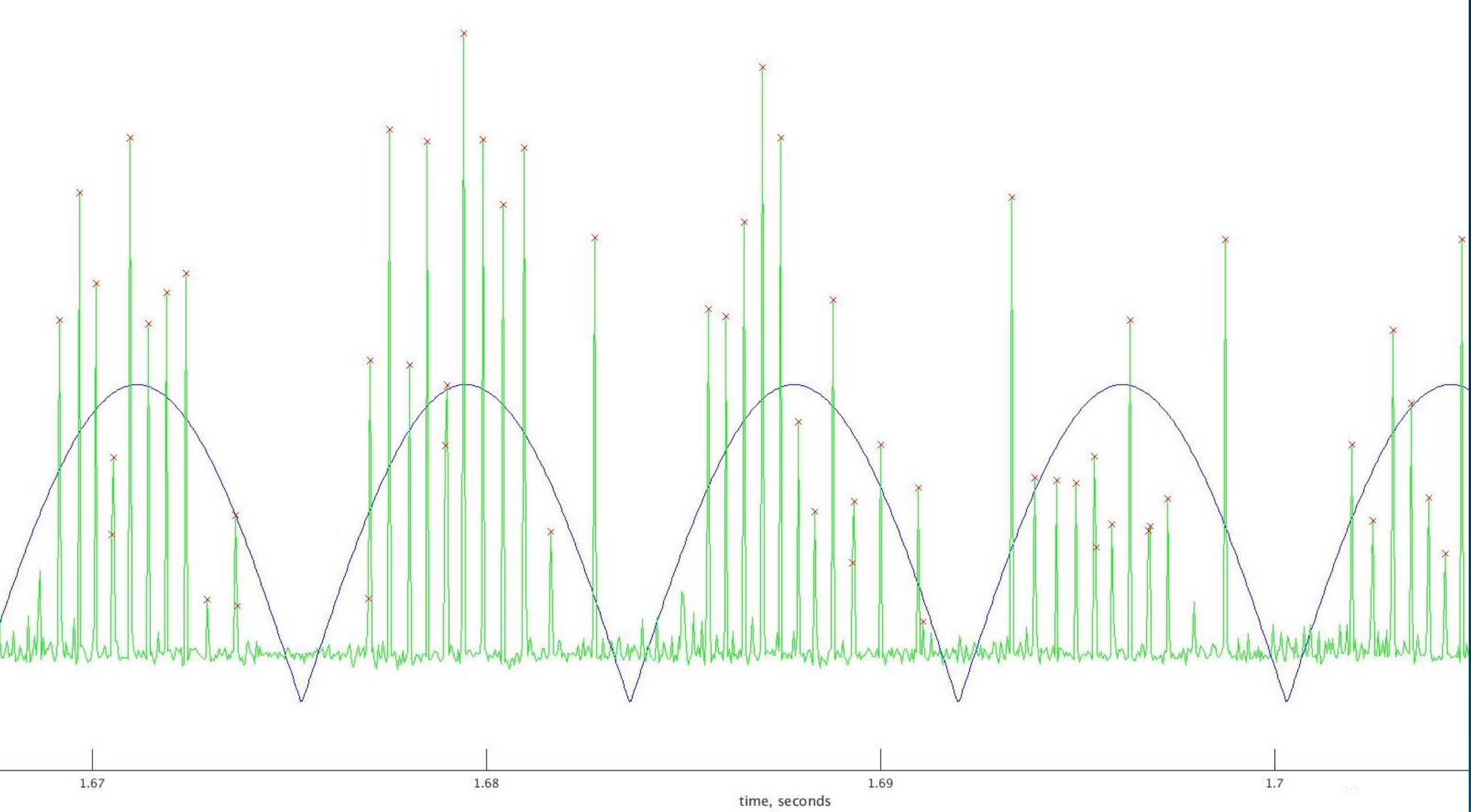


# FPGA

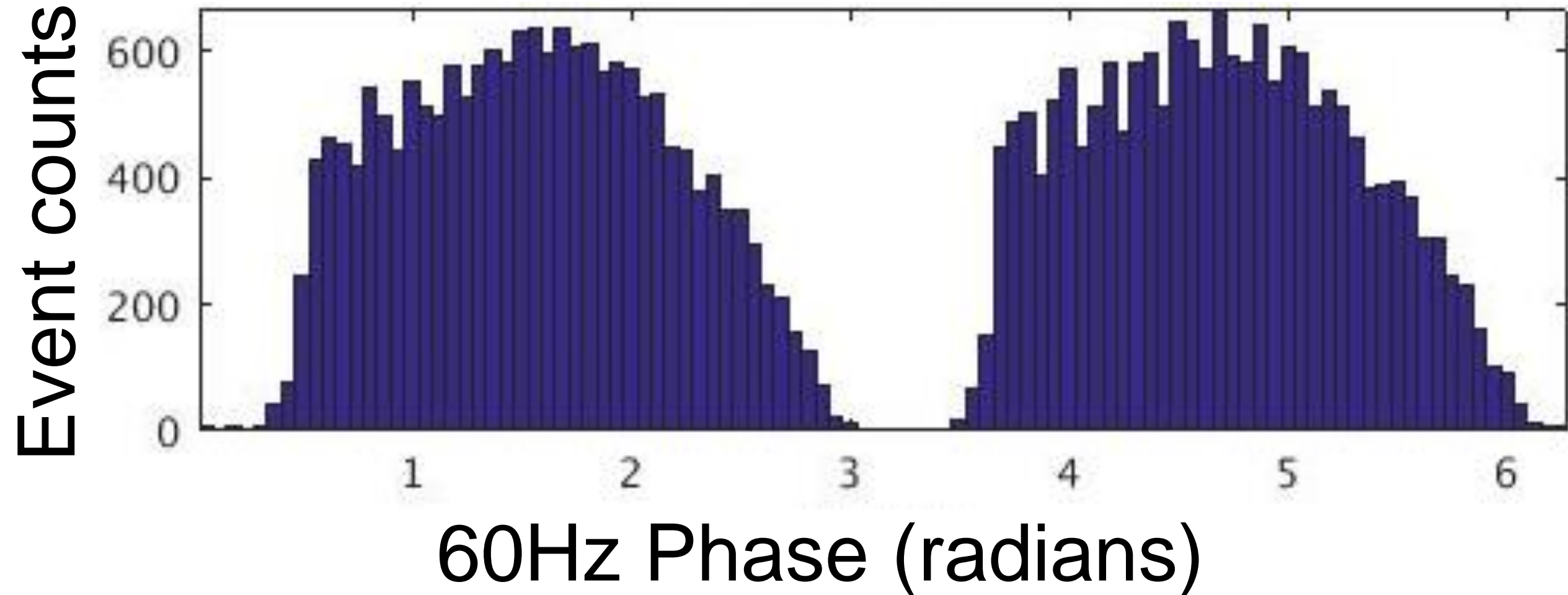


**OVRO-LWA antennas, colorized by attendant FPGA**





# Power line RFI histogram vs voltage phase





# Next Steps

- Voltage time of arrival solutions only good to RMS 3 samples
  - Improvement = better localization
- 1.6M events; probably 15 CRs in data
  - How to convince self that detection is actually cosmic ray?
- “More (or better) data trumps better algorithms”
  - Work on fixing bad RFI sources

# Takeaways

- RFI is a problem.
  - Planes preclude a fully automatic pipeline
  - Could we run from much of our RFI by tracking the AC voltage?
  - This tool is very powerful for localizing RFI!
- Need a system which is matched to the CR detection task
  - Bring all the voltages home
  - Ideally, process all the voltages in the same location
- I am very close!